

Multi-Year Observations of Ultra- Relativistic Electrons with the Van Allen Probes Mission

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Discovering Earth's radiation belts

Daniel Baker and
Mikhail Panasyuk

Six decades after the belts' discovery in 1958, scientists are still finding new mysteries.

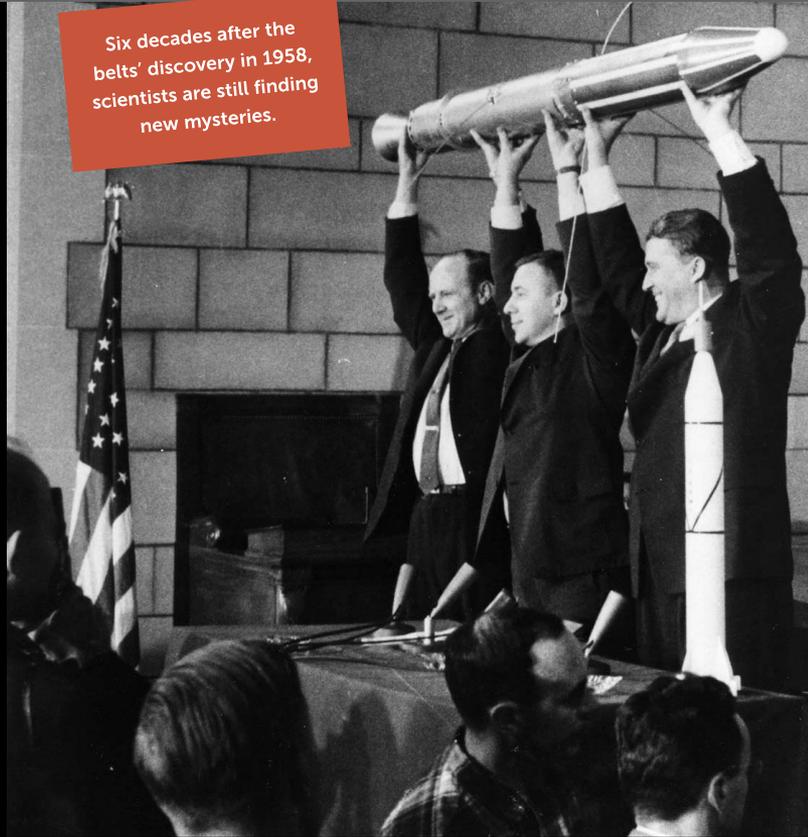
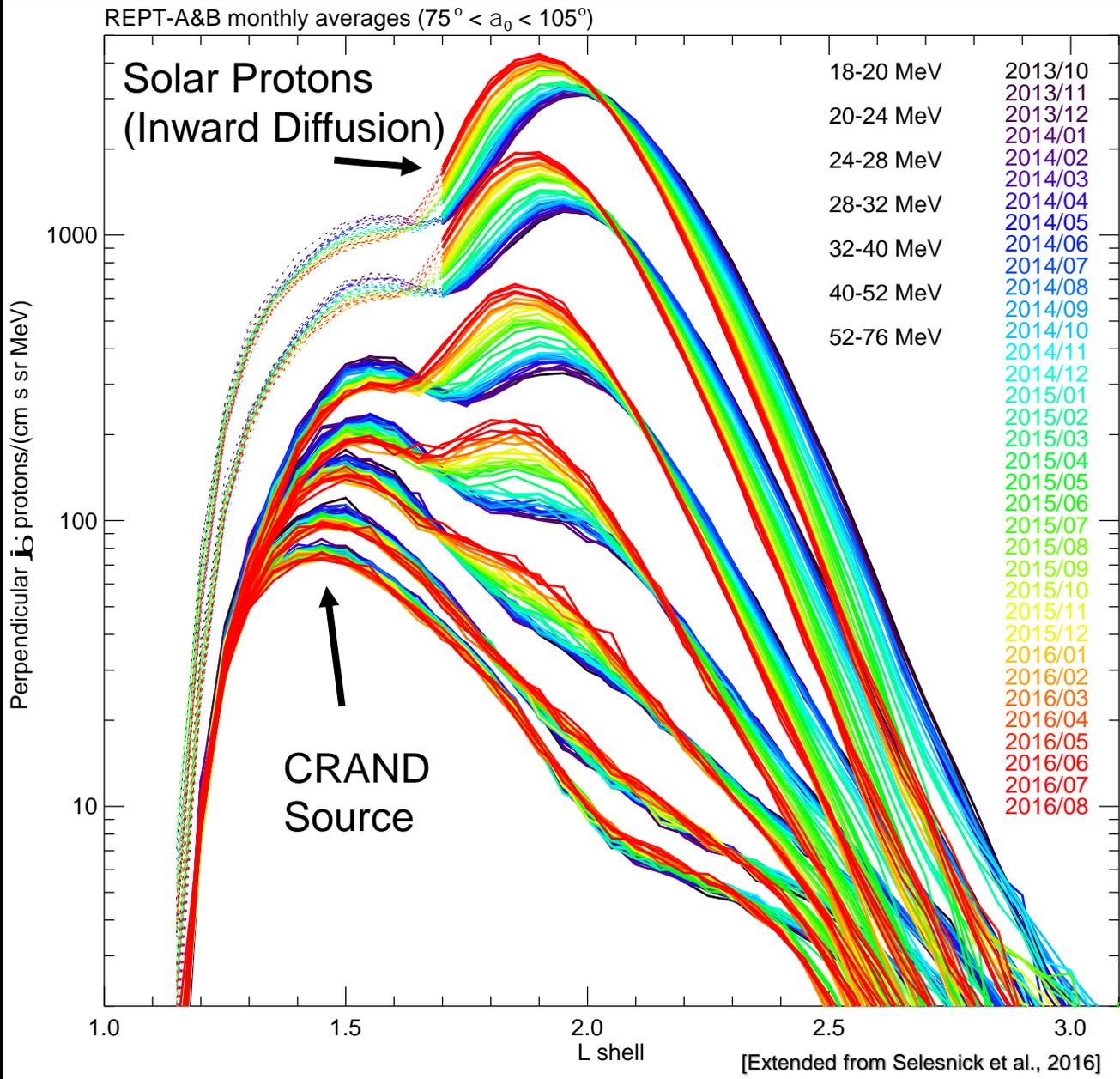


Photo taken at
U.S. National
Academy Building
(31 January 1958)

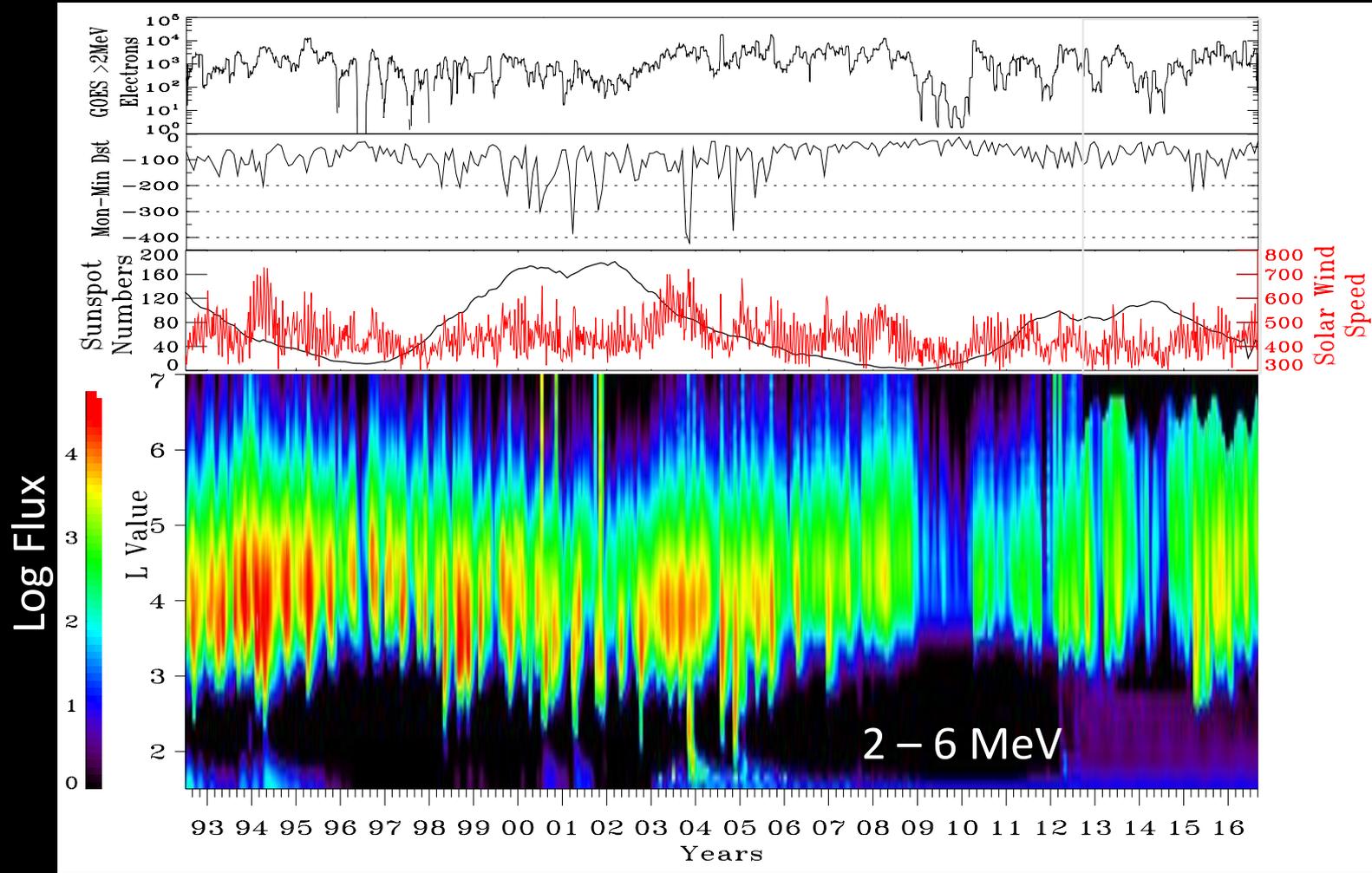
Explorer 1
Launch

2 PHYSICS TODAY | DECEMBER 2017

REPT Monthly Average Proton Flux (October 2013 – August 2016)

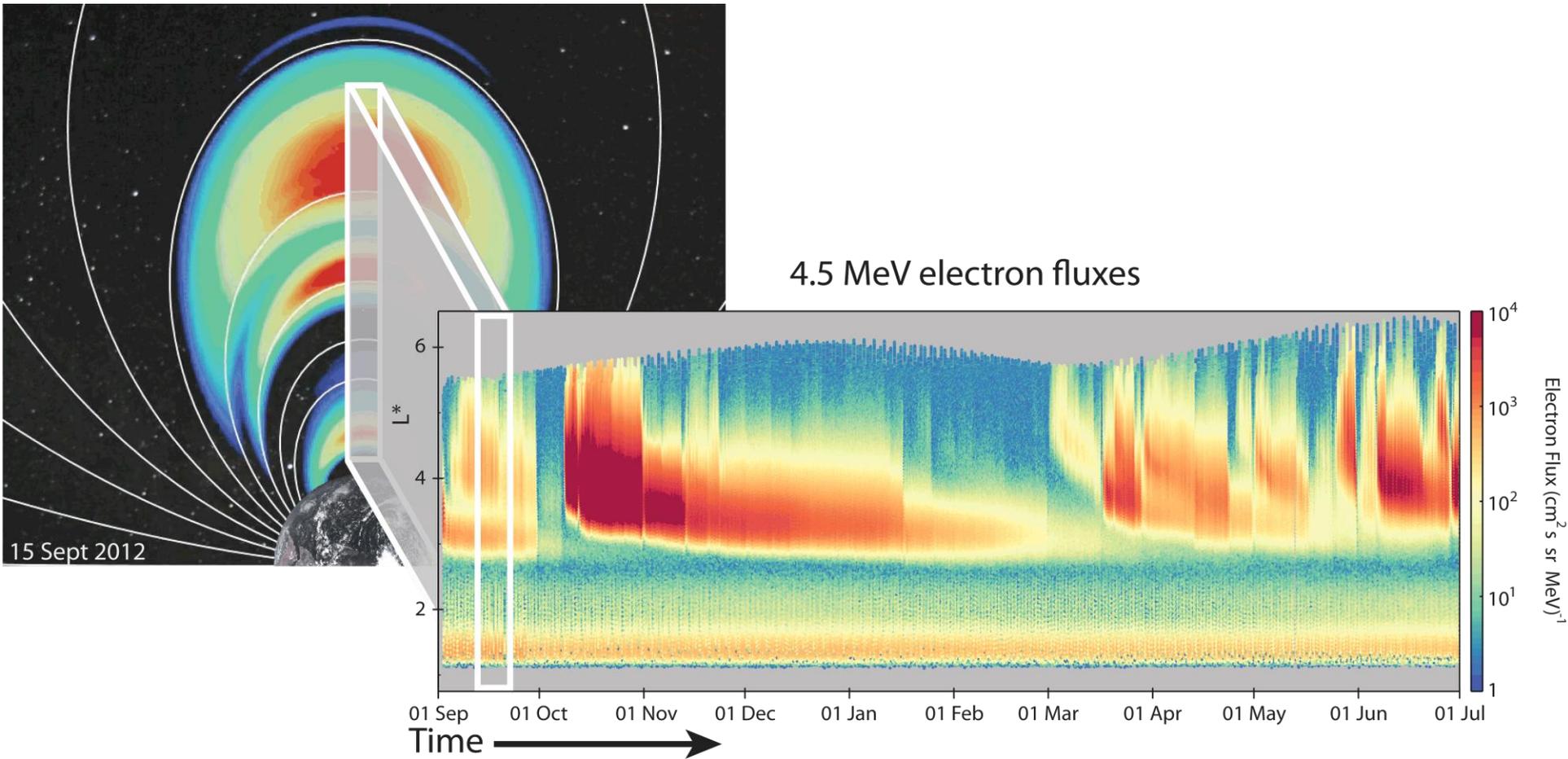


SAMPEX

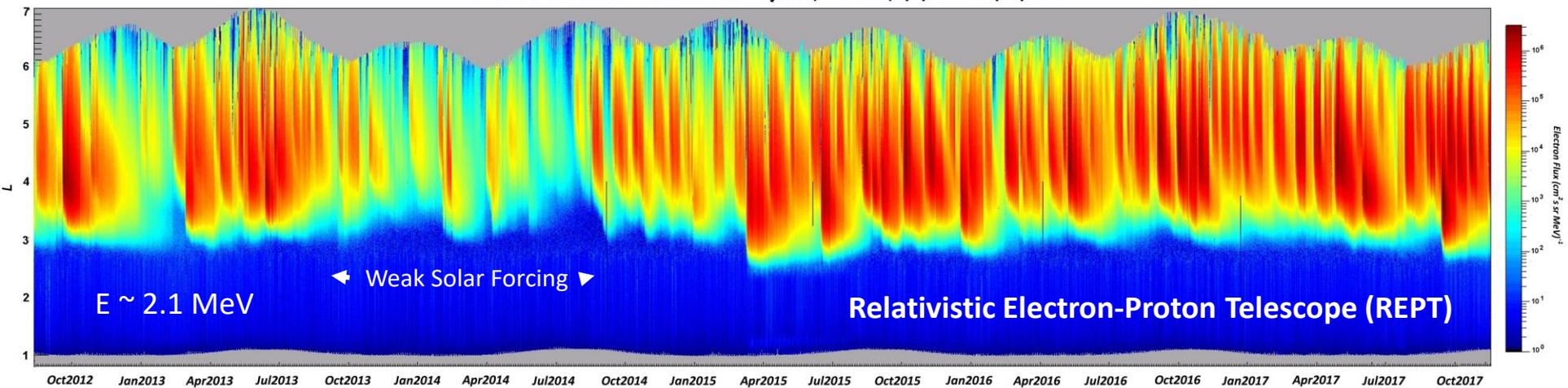


REPT Observations:

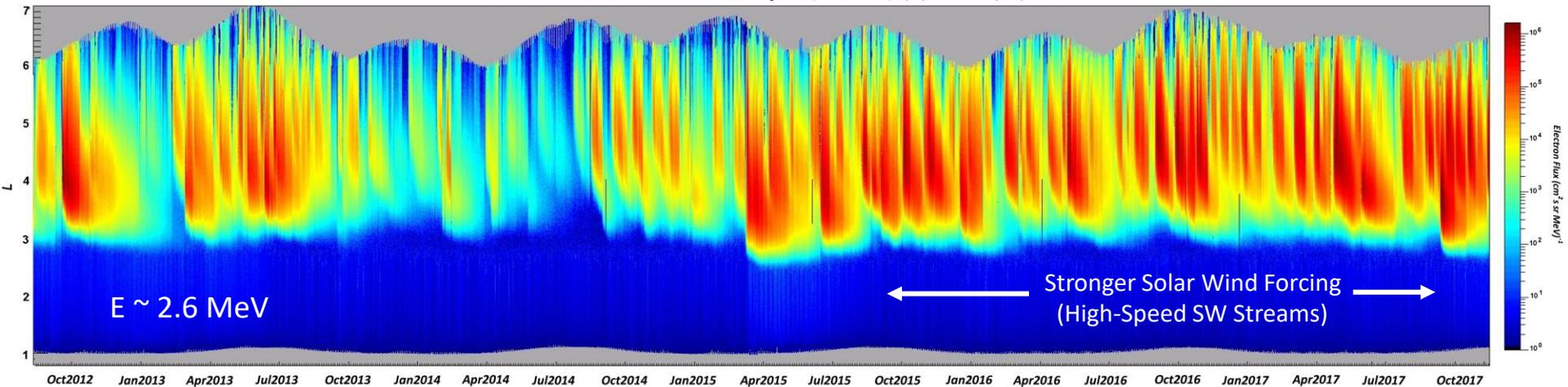
Acceleration, Remanence, and Sudden Loss



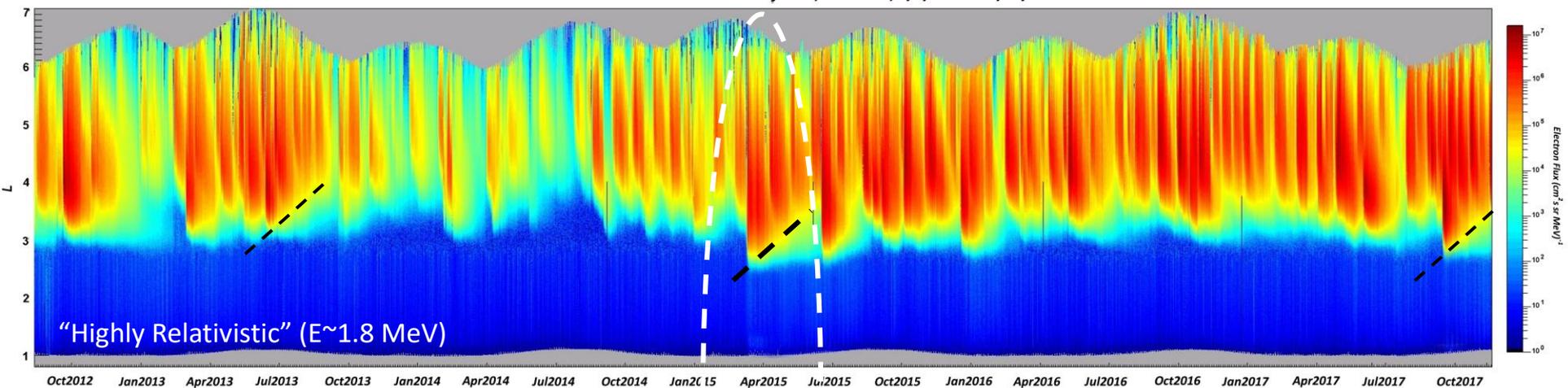
RBSP ECT-REPT A & B 2.1 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



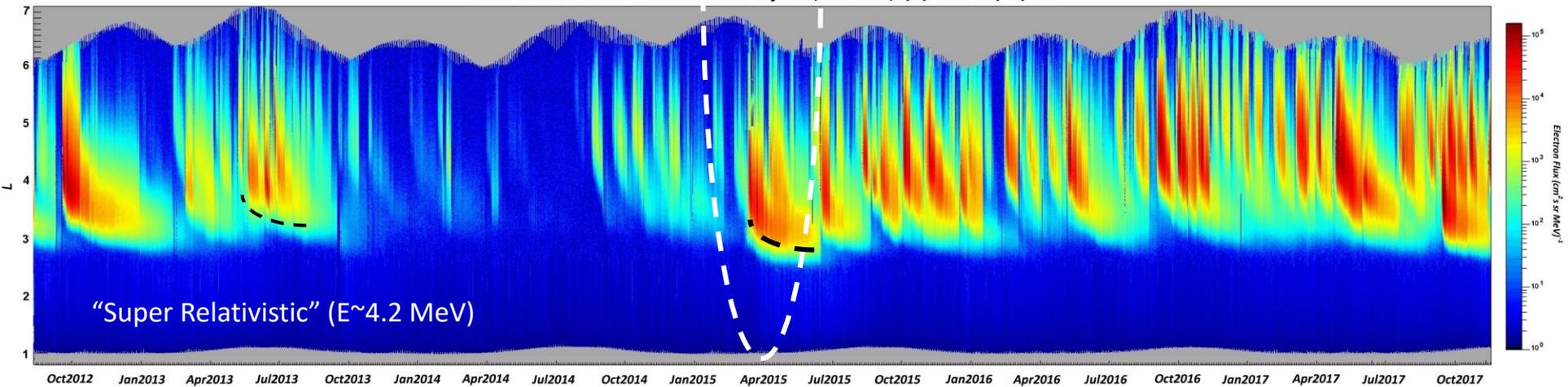
RBSP ECT-REPT A & B 2.6 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



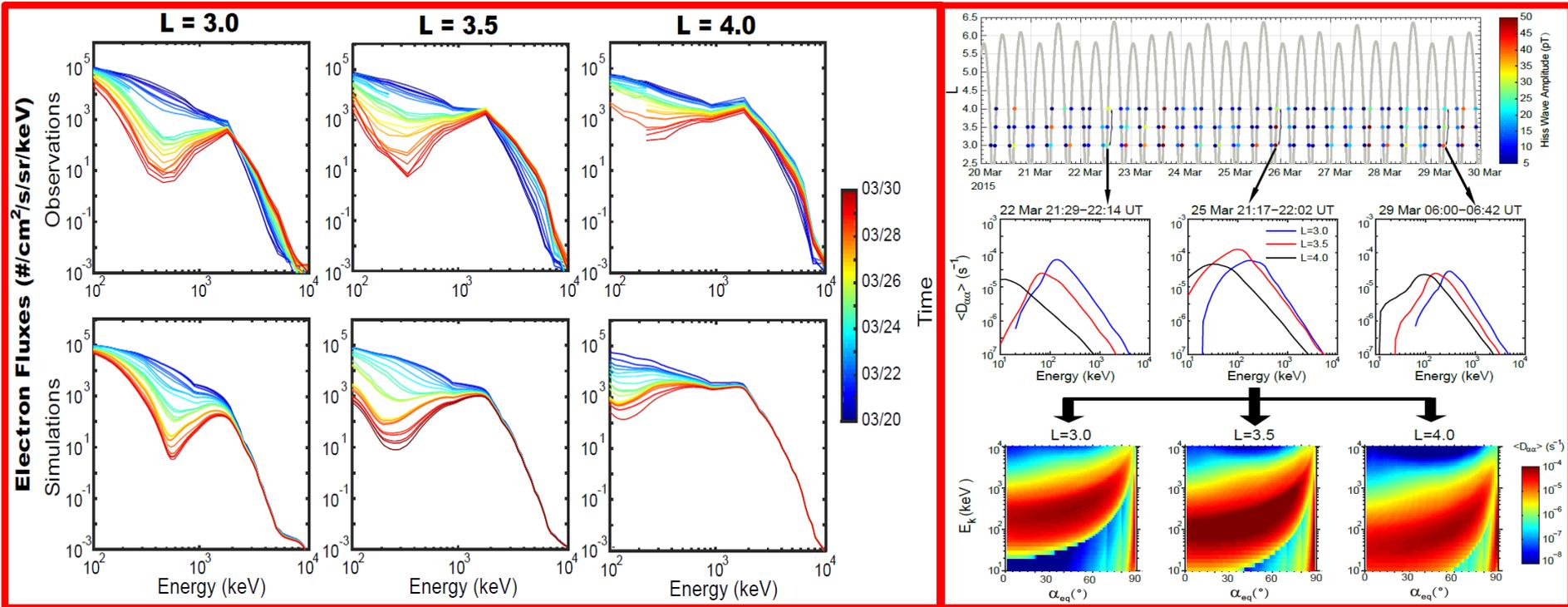
RBSP ECT-REPT A & B 1.8 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



RBSP ECT-REPT A & B 4.2 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



Radiation Belt Electron Bump-On-Tail Distribution and Its Origin from Hiss Induced Scattering



[Courtesy of B. Ni, Wuhan University]

2015/03/20 – 03/30 [Comparisons](#)

Top: BOT observations

Bottom: BOT simulations

[Hiss Induced Scattering](#)

2015/03/20 – 03/30

Top:

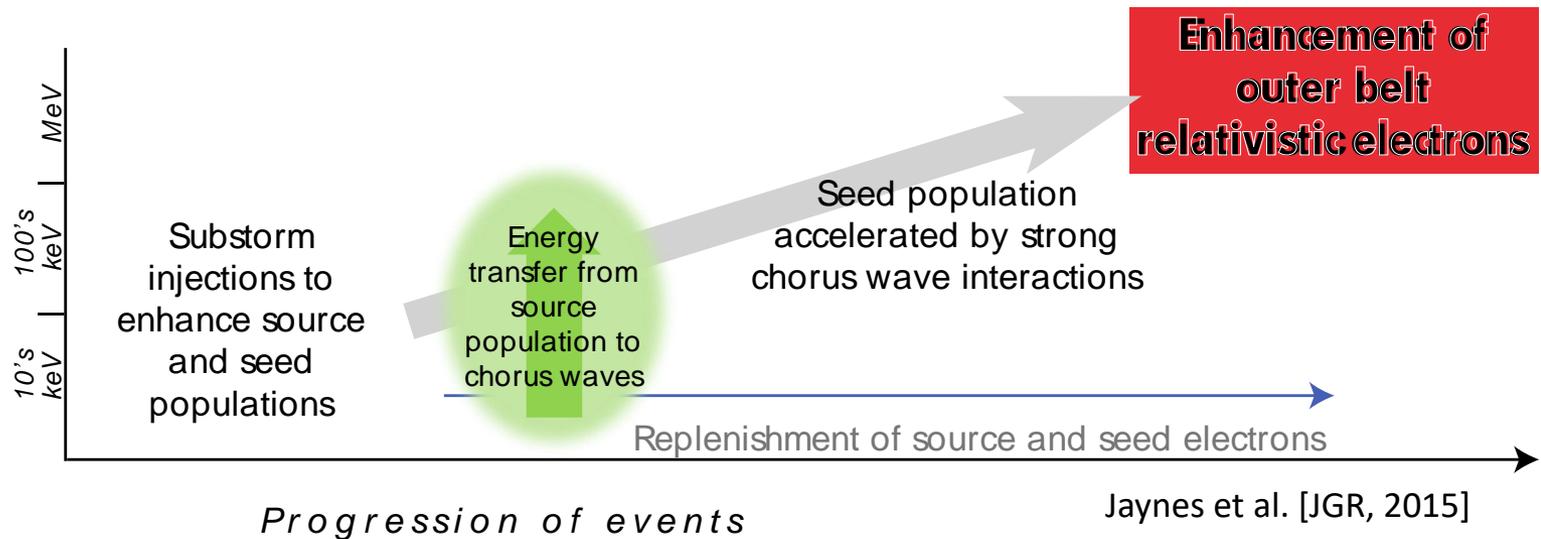
10-day Hiss wave data

Middle: $\langle D_{\alpha\alpha} \rangle$ at loss cone at three time intervals

Bottom: 2-D plot of $\langle D_{\alpha\alpha} \rangle$ at one time interval

Work led by Hong Zhao and B.Ni

Electron Acceleration Sequence in the Inner Magnetosphere

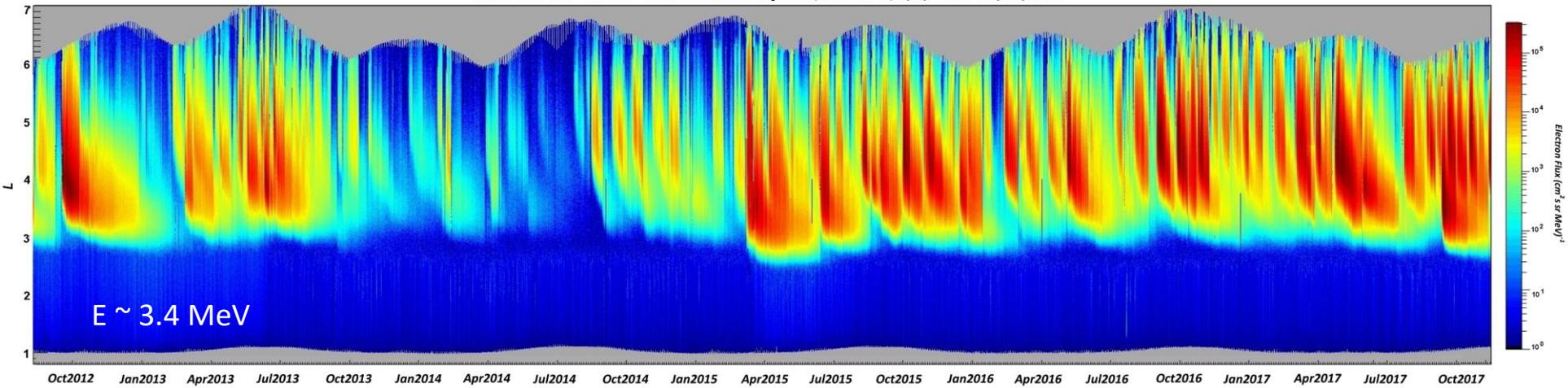


Source: 10's of keV
Seed: 100's of keV
Rad Belt: > 1 MeV

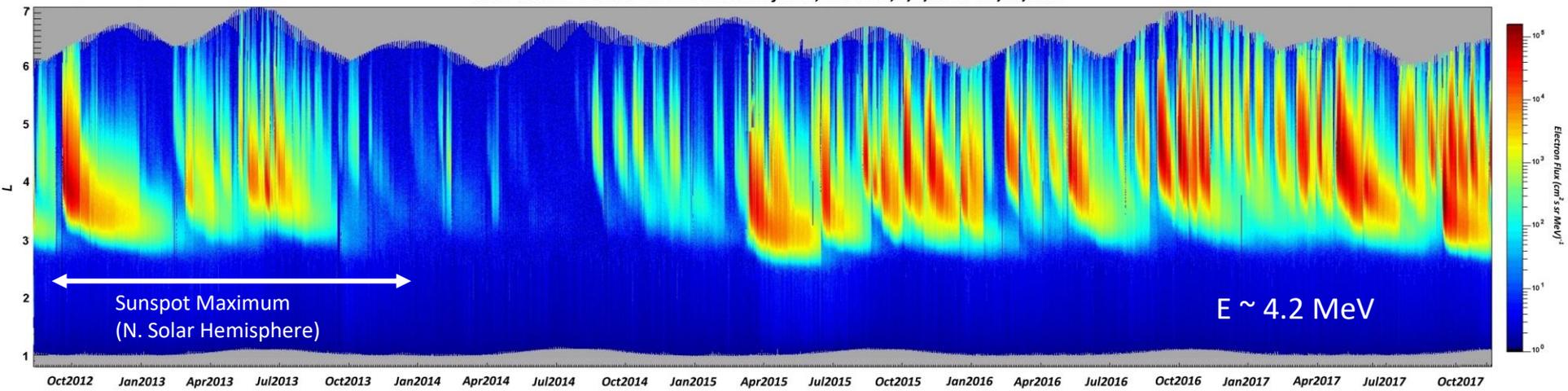
Crucial Role for
Magnetospheric
Substorms

Jaynes et al. (JGR, 2015)

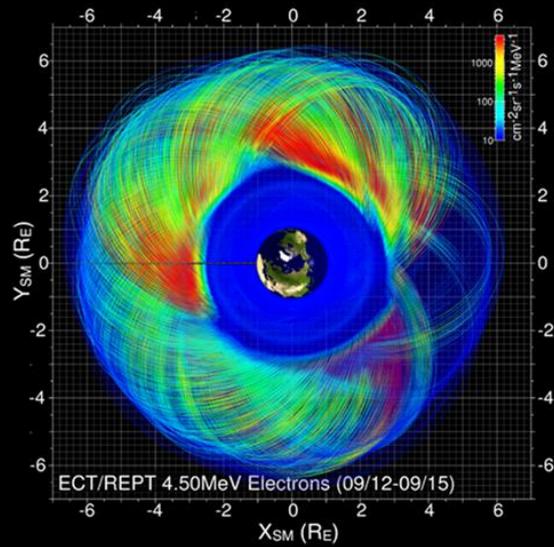
RBSP ECT-REPT A & B 3.4 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



RBSP ECT-REPT A & B 4.2 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017

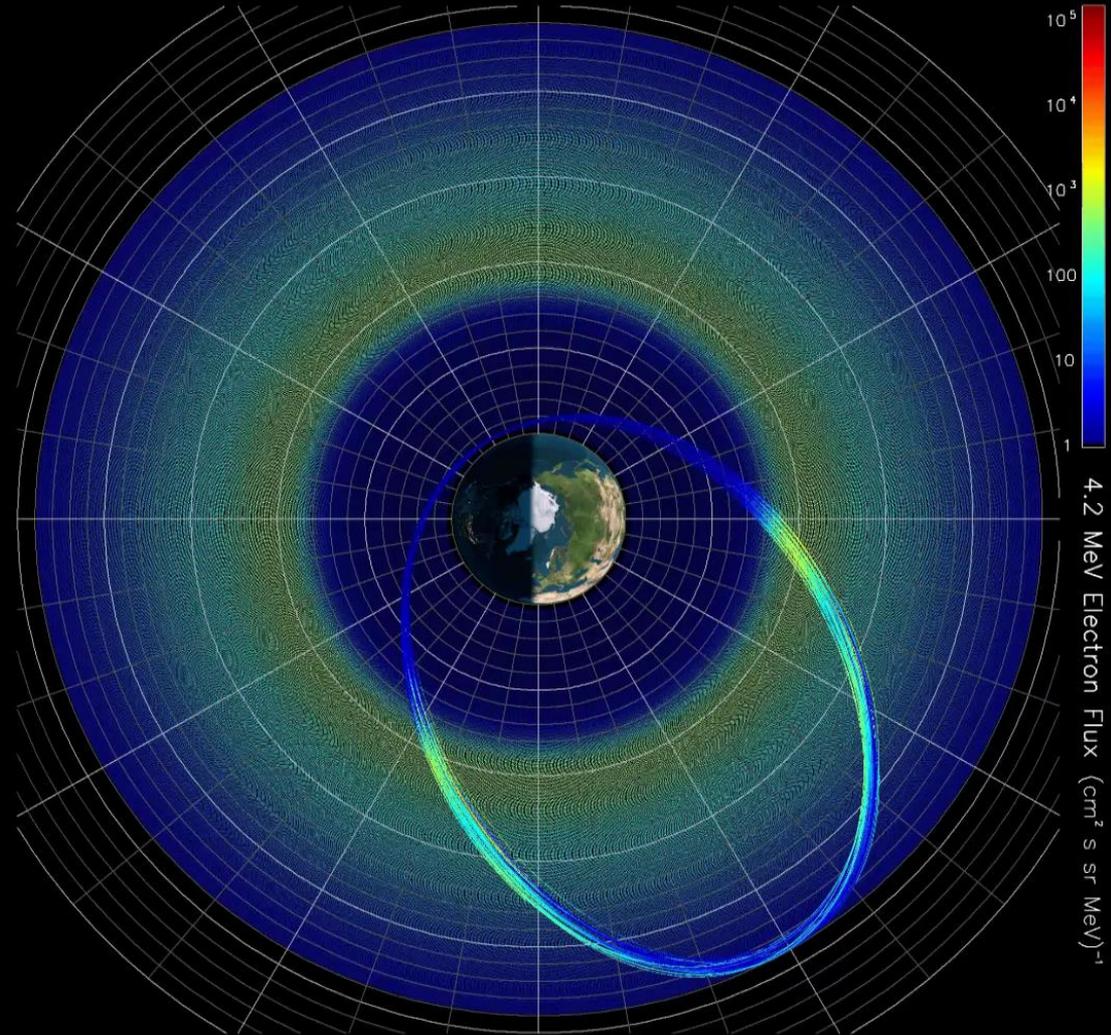


REPT – The Movie

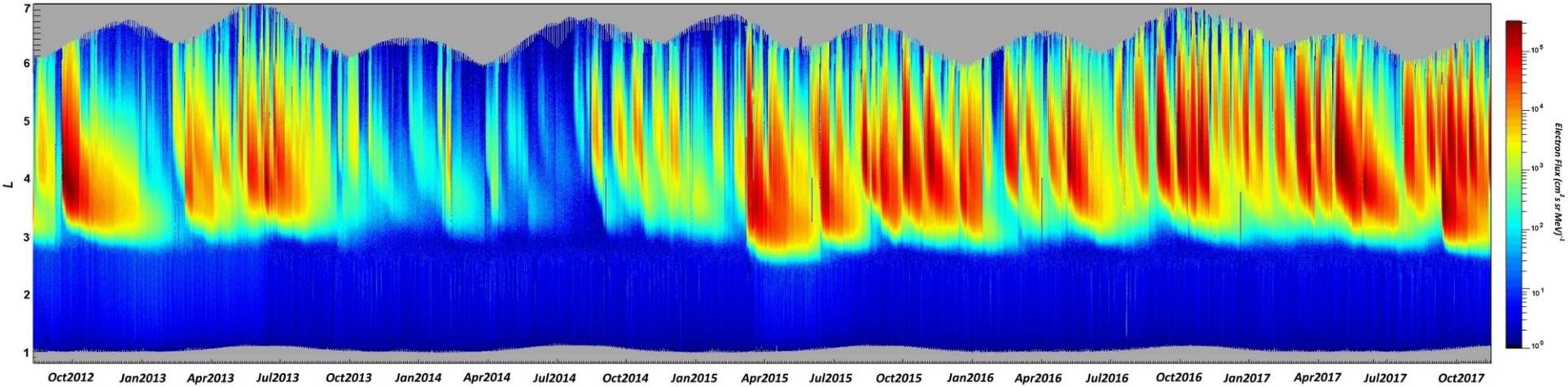


Van Allen Probes ECT-REPT A & B

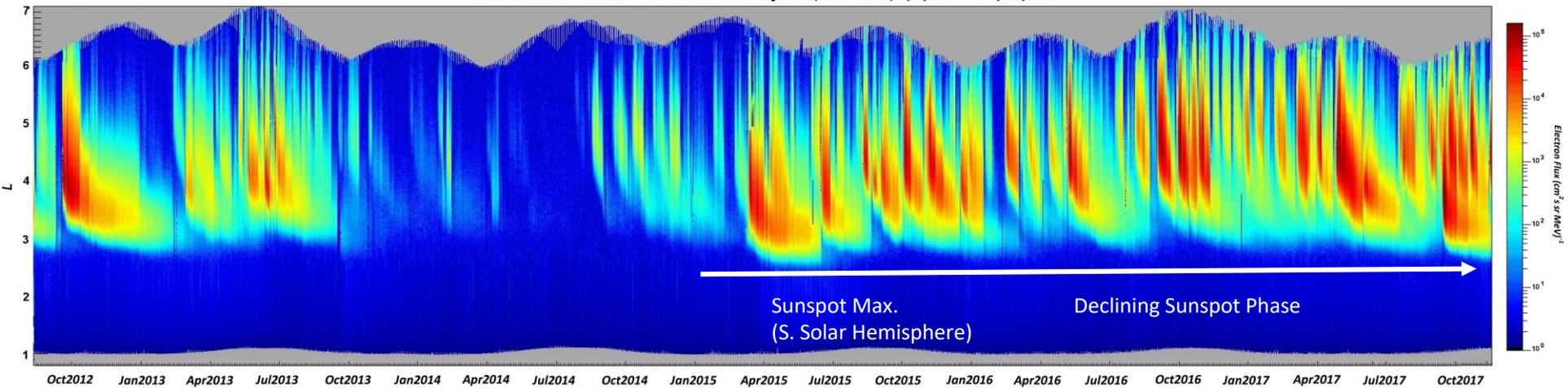
2012-09-08



RBSP ECT-REPT A & B 3.4 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



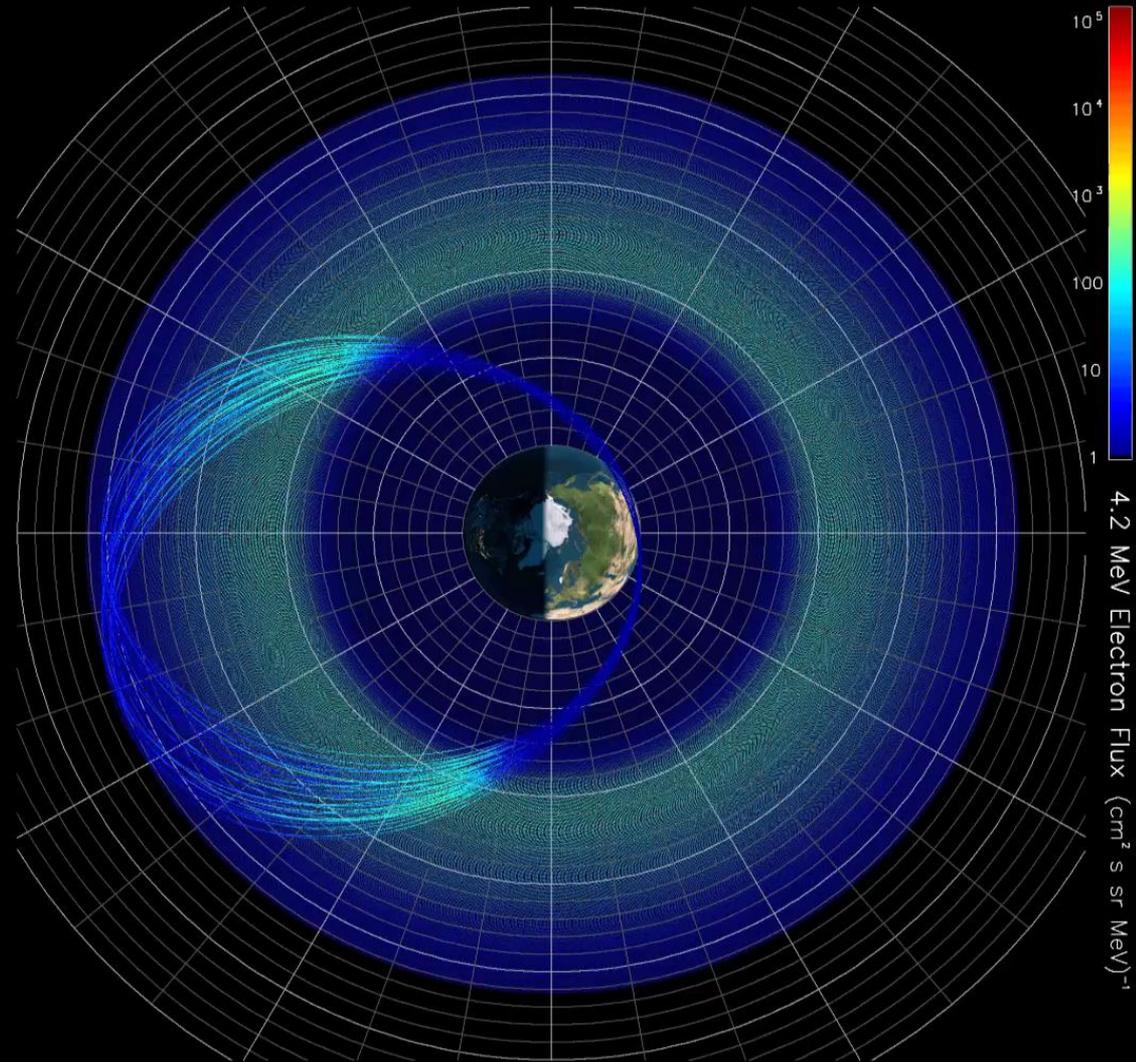
RBSP ECT-REPT A & B 4.2 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



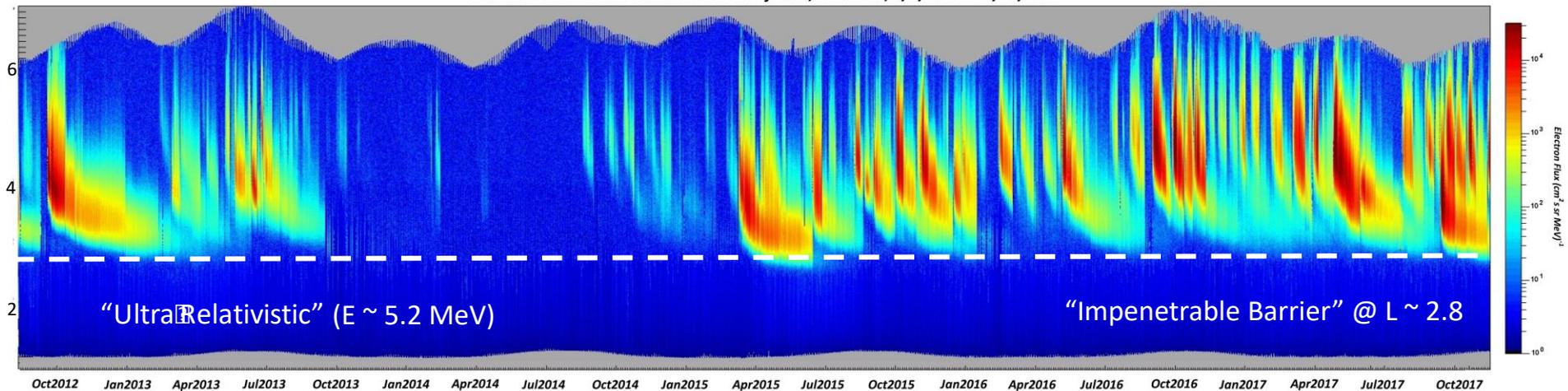
REPT – 2015 to 2017

Van Allen Probes ECT-REPT A & B

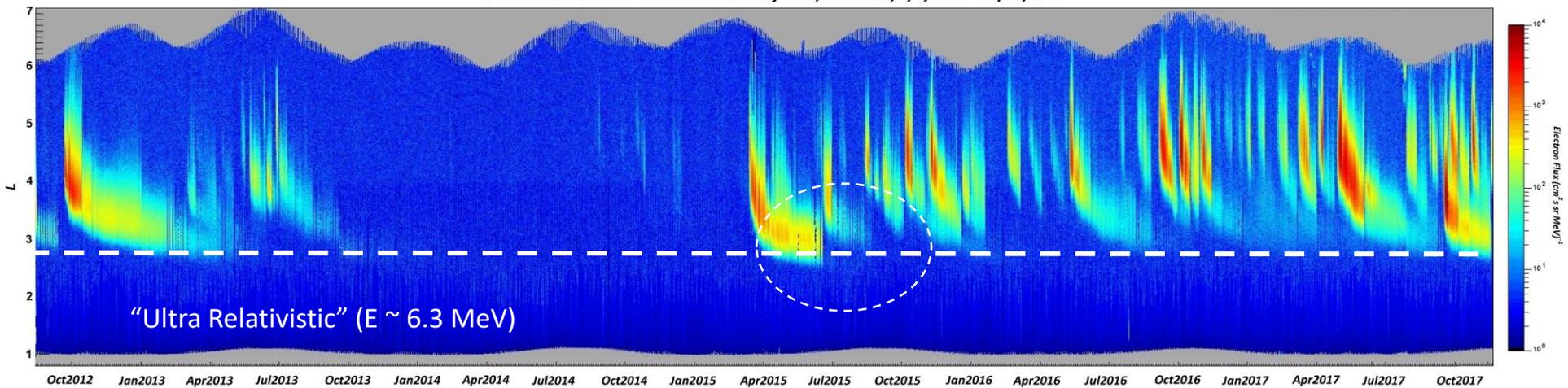
2015-01-01



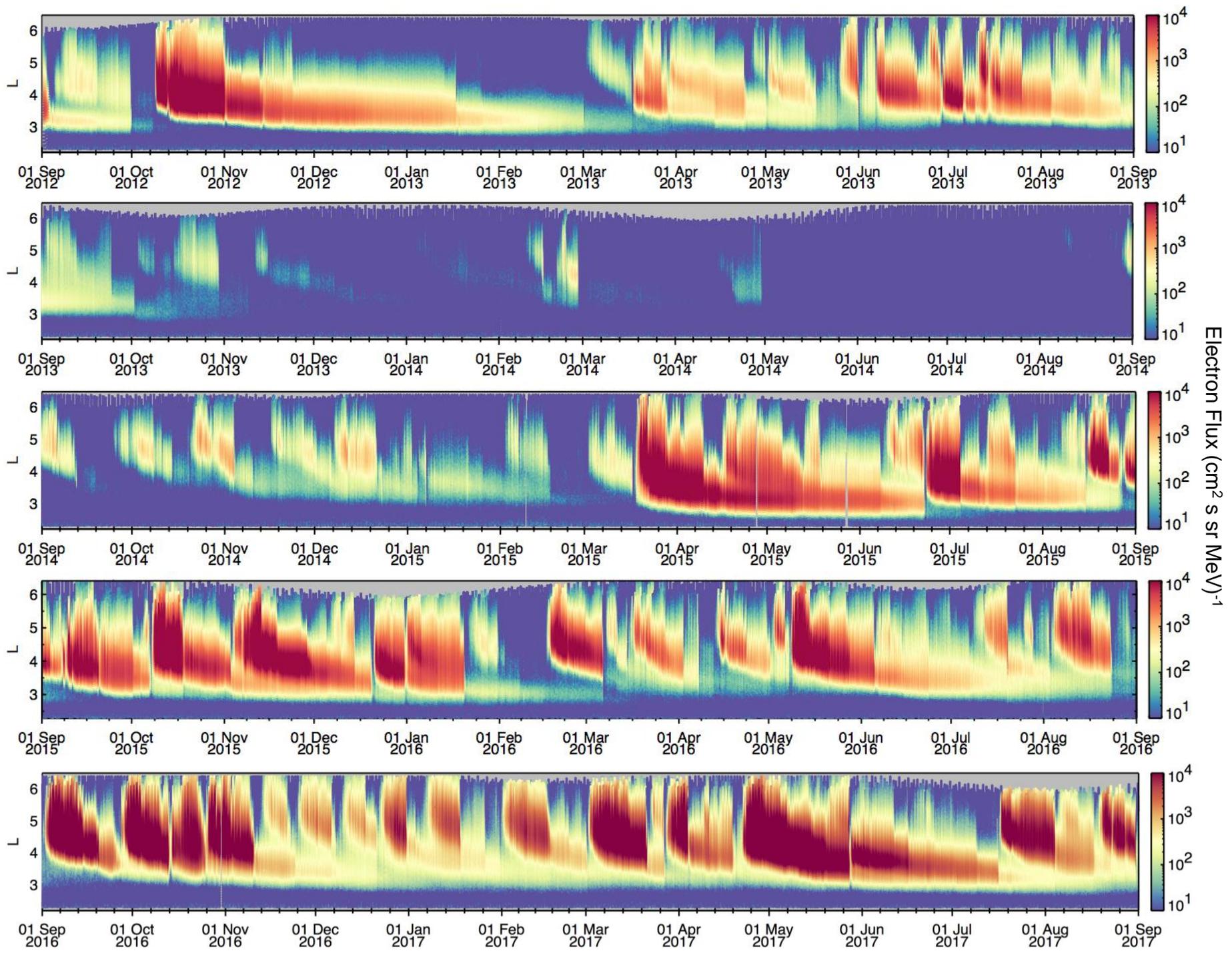
RBSP ECT-REPT A & B 5.2 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



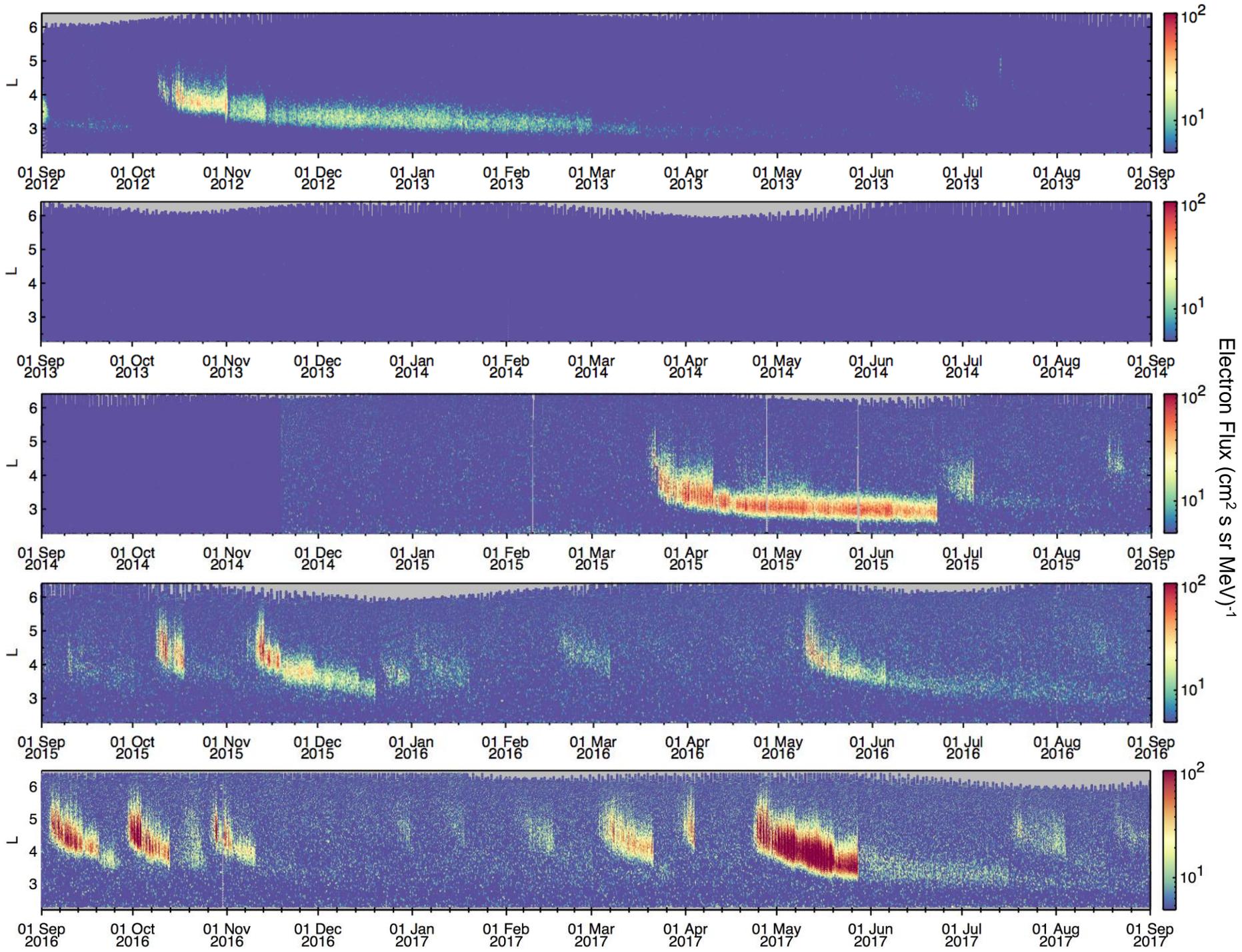
RBSP ECT-REPT A & B 6.3 MeV Electron fluxes, L vs Time, 9/1/2012 - 11/10/2017



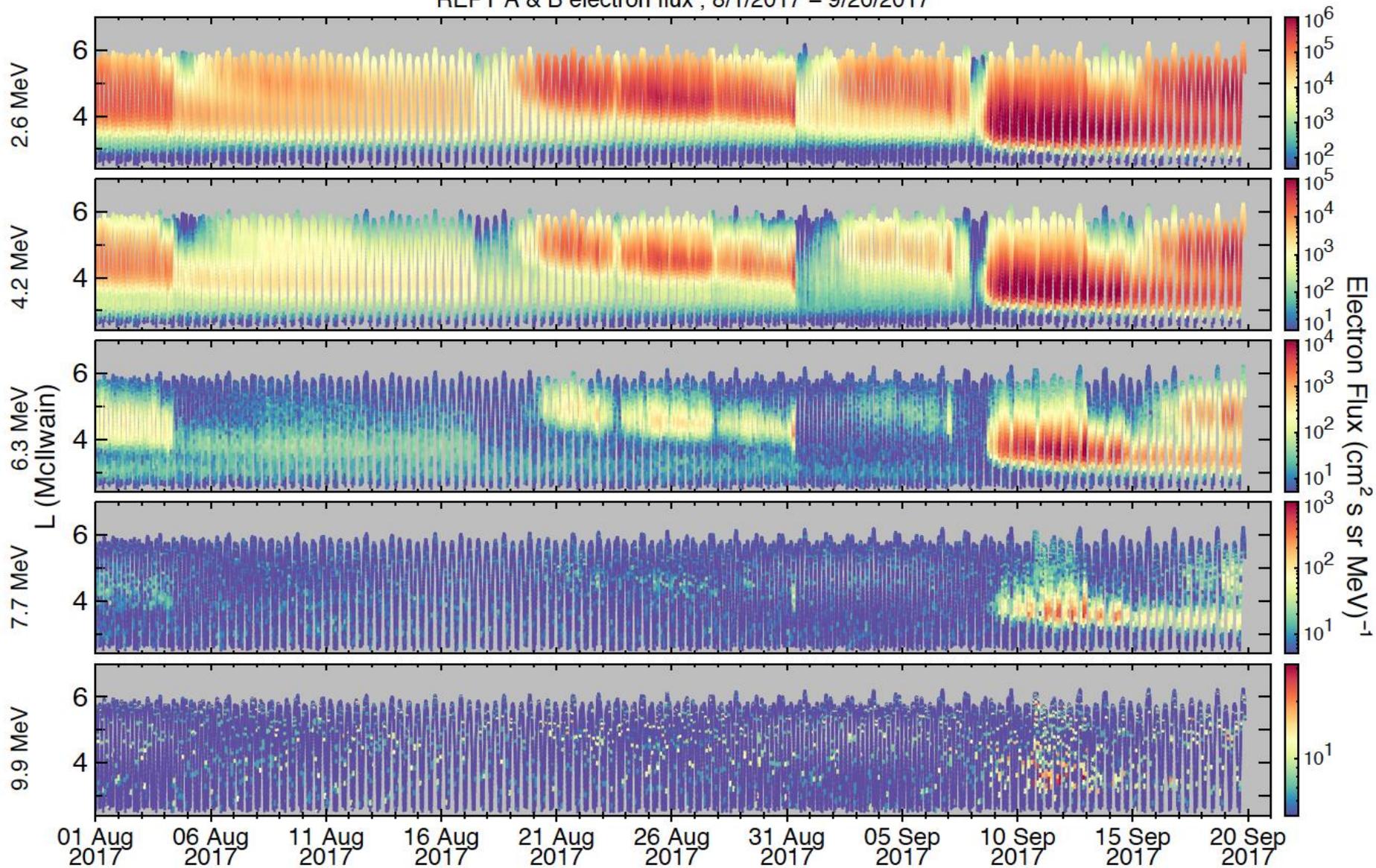
REPT A & B 4.2 MeV Electrons



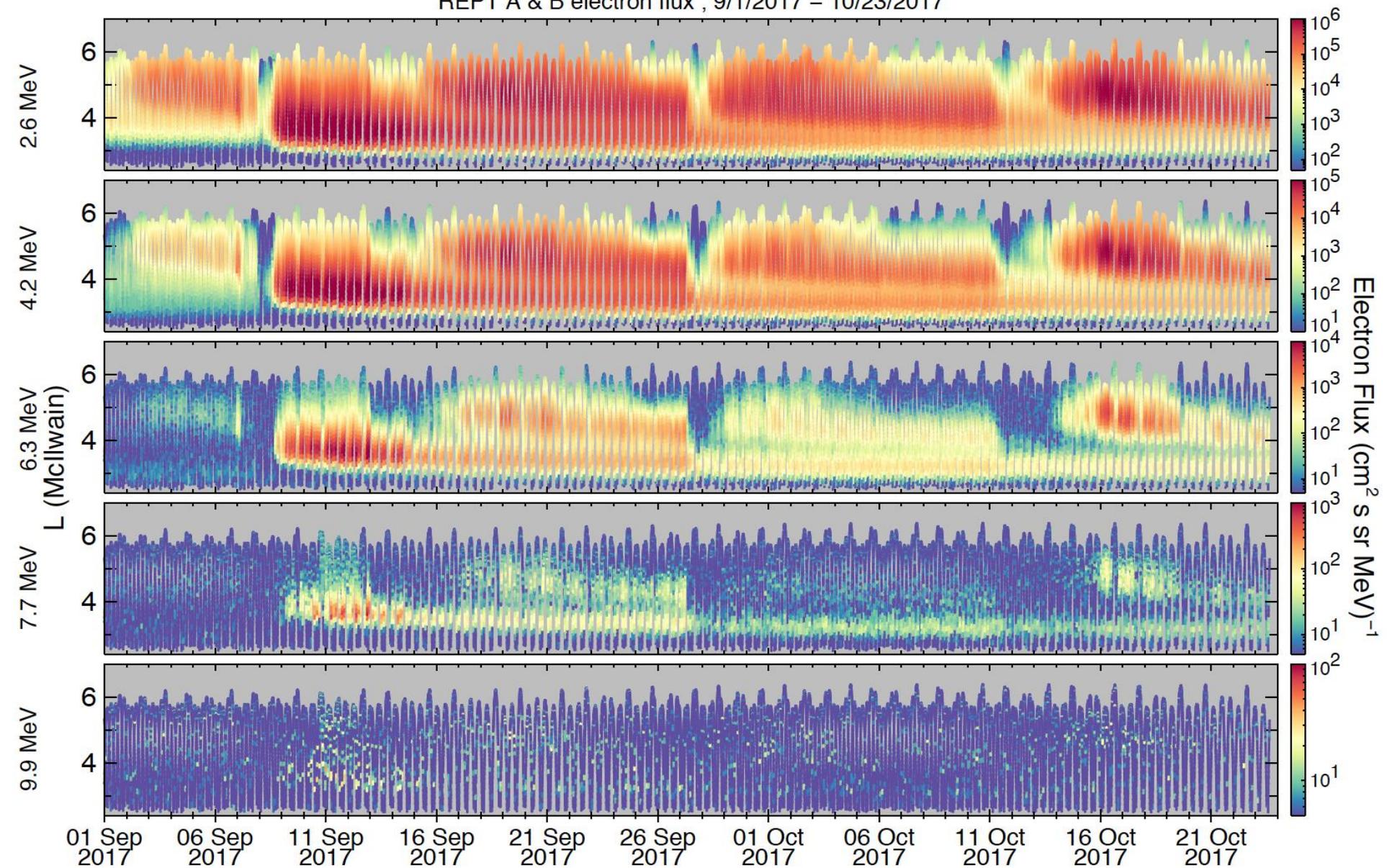
REPTA & BIT7 7 MeV Electrons



REPT A & B electron flux , 8/1/2017 – 9/20/2017



REPT A & B electron flux , 9/1/2017 – 10/23/2017

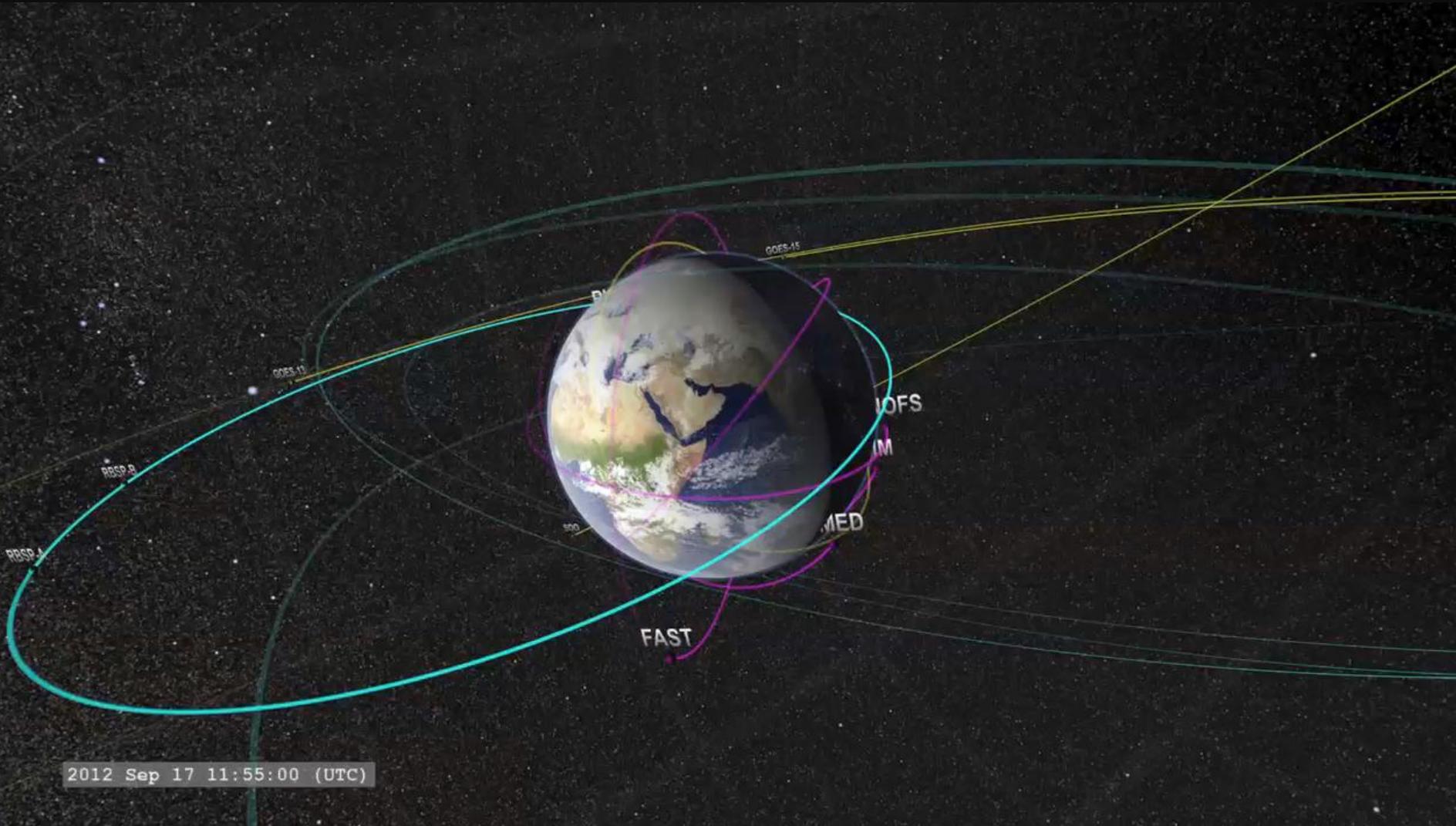
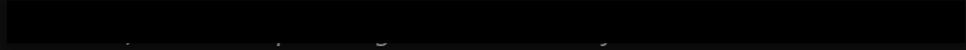


Conclusions

- Results from the Van Allen Probes mission demonstrate remarkable, previously unobserved features about radiation belt structure, acceleration, transport, and rapid loss.
- Long-term observations reveal distinctive behavior: Multi-belt structure and impenetrable barrier to inward penetration of ultra-relativistic electrons at $L \sim 2.8$: No cases of high fluxes of $E > 1.5$ MeV electrons inside of $L \sim 2.5$ in over five years of measurements.
- Van Allen Probes data clearly show there are extended periods of gradual change in the (super- and ultra-) relativistic electron populations punctuated by abrupt losses and rapid subsequent acceleration.
- REPT instruments show that ultra-relativistic electrons were low around 2014 sunspot max and have now been increasing dramatically due to strong solar wind streams in declining sunspot phase (southern solar hemisphere).

Questions?

The Geospace Observatory Fleet



REPT A & B 1.8 MeV Electrons

