The storm that wasn't: A look at multiple loss processes occurring simultaneously and how they interact with each other

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March 8th 2018 Chapman Conference: Particle Dynamics in the Earth's Magnetosphere Cascais Portugal

January 7th 2014 the CME and iCME-shock propagation



Möstl, C. et al. (2015), Strong coronal channeling and interplanetary evolution of a solar storm up to Earth and Mars, Nature Communications, doi: 10.1038/ncomms8135

January 9th 2014 Arrival of the ICME-Shock at Earth



9 January 2014 UT

Halford, A. J., et al. (2015), BARREL observations of an ICME-shock impact with the magnetosphere and the resultant radiation belt electron loss. J. Geophys. Res. Space Physics, doi: <u>10.1002/2014JA020873</u>.

January 9th 2014 array of observations



Halford, A. J., et al. (2015), BARREL observations of an ICME-shock impact with the magnetosphere and the resultant radiation belt electron loss. J. Geophys. Res. Space Physics, doi: <u>10.1002/2014JA020873</u>.

January 9th 2014 overview of shock impact on the magnetosphere:

CME

ICME-Shock

NASA/Goddard Space Flight Center 12:00

315

B

January 9th 2014 overview of shock impact on the magnetosphere: E-field impulse

CNE

315

В

З

12:00

ICME-Shock

Not to scale

January 9th 2014 overview of shock impact on the magnetosphere: Generation of cyclotron resonant waves

CNE

ICME-Shock

Not to scale

G15

B

з

12:00

January 9th 2014 observations of inferred electron precipitation: BARREL observations



Halford, A. J., et al. (2015), BARREL observations of an ICME-shock impact with the magnetosphere and the resultant radiation belt electron loss. J. Geophys. Res. Space Physics, doi: <u>10.1002/2014JA020873</u>.

January 9th 2014 observations of Electric field impulse: RBSP observations



Halford, A. J., et al. (2015), BARREL observations of an ICME-shock impact with the magnetosphere and the resultant radiation belt electron loss. J. Geophys. Res. Space Physics, doi: <u>10.1002/2014JA020873</u>.

January 9th 2014 observations of Electric field impulse: Expected effect on the local particle population

Conservation of the first and second adiabatic invariants

$$L_o = \left[\frac{R_E B_o L_f^2}{2E\delta t L_f^2 + R_E B_o}\right]^{1/2}$$

e.g. Wygant et al 1994

$$\sin\alpha_{eq_{f}} = \left[\frac{-L_{f}^{1/2}\cos^{2}\alpha_{eq_{o}}}{2L_{o}^{1/2}\sin\alpha_{eq_{o}}}\right] + \frac{1}{2}\left[\frac{L_{f}\cos^{4}\alpha_{eq_{o}}}{L_{o}\sin^{2}\alpha_{eq_{o}}} + 4\right]^{1/2}$$

e.g. Shultz and Lanzerotti 1974

$$L_f = 5.8, L_o = 6.8 \Rightarrow \alpha_{eq_o} < \alpha_{eq_f}$$



Halford, A. J., et al. (2015), BARREL observations of an ICME-shock impact with the magnetosphere and the resultant radiation belt electron loss. J. Geophys. Res. Space Physics, doi: <u>10.1002/2014JA020873</u>. January 9th 2014 observations of Electric field impulse: Expected effect on the local particle population



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January 9th 2014 observations of Electric field impulse: Expected effect on the local particle population



Halford, A. J., et al. Submitted to JGR

January 9th 2014 observations of Electric field impulse: Expected generation of an EMIC wave



$$S = \frac{\mu}{V_g} = \frac{\sum_s \frac{\eta_{sw} \sqrt{\pi}}{M_s^2 \alpha_{||,s}} \Big[(A_s + 1)(1 - M_s X) - 1 \Big] \times exp \Bigg[\frac{\frac{-\eta_{s,w}}{M_s} \frac{(M_s X - 1)^2}{\beta_{sw} X^2}}{\frac{(1+\delta)}{(1-X)} + \sum_j \frac{(\eta_{jw} + \eta_{jc})M_j}{1 - M_j X}} \Bigg]}{2X^2 \Big[\frac{1+\delta}{1-X} + \sum_j \Big(\eta_{jw} + \eta_{jc} \frac{M_j}{1 - M_j X} \Big) \Big]}$$

e.g. Halford 2012, Fraser et al 1989. Gomberoff and Neira 1983, and Kozyra et al 1984

January 9th 2014 observations of EMIC wave: Expected effect on the local particle population



Halford, A. J., et al. in prep

January 9th 2014 observations of Electric field impulse: Expected generation of an Chorus wave



Halford, A. J., et al. Submitted to JGR

January 9th 2014 observations of Chorus wave: Expected effect on the local particle population

$$D_{\alpha,\alpha} = \frac{\pi \Omega_{\sigma}^2}{2\upsilon |\Omega_e|} \frac{1}{(E+1)^2} \sum_s \sum_j \frac{R\left(1 - \frac{x\cos\alpha}{y\beta}\right)^2 |F(x,y)|}{\delta x |(\beta\cos\alpha - F(x,y))|} e^{-\left(\frac{x-x_m}{\delta x}\right)^2}$$

Summers et al 2005, 2007



Halford, A. J., et al. (2015), BARREL observations of an ICME-shock impact with the magnetosphere and the resultant radiation belt electron loss. J. Geophys. Res. Space Physics, doi: <u>10.1002/2014JA020873</u>.

January 9th 2014 observations of Chorus wave: Expected effect on the local particle population



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January 9th 2014 observations of Chorus wave: Expected effect on the local particle population



Normalized Flux

Halford, A. J., et al. in prep

January 9th 2014 observations of Electric field impulse: Expected generation of a Hiss wave



Halford, A. J., et al. Submitted to JGR

The Solar Wind Shock Impact: Particles

- 1. Pushed the particles earthward by 1 RE
- 2. Particles within 0.5 deg of the loss cone are lost to the atmosphere
- 3. This may be the cause for the temporal structure and MLT difference between the payloads

12:00

The Solar Wind Shock Impact: waves

- 1. Generated an increase in temperature anisotropy
- 2. Resulted in growth of EMIC and chorus waves simultaneously at RBSP-B outside of the plasmasphere
- 3. Resulted in growth of Hiss at RBSP-A inside the plasmasphere.

12:00

3

The EMIC wave Impact: Particles

- 1. Caused the loss of MeV electrons into the upper atmosphere
- 2. However, very little loss at these energies was observed as inferred by the balloons

12:00

E)

3

The Chorus wave Impact: Particles

- 1. Caused the loss of eV 100s keV electrons into the upper atmosphere.
- 2. However, very little loss at these energies was observed as inferred by the balloons

12:00

B

3

The Hiss wave Impact: Particles

- 1. Time scales appear to have been too small to be able to infer significant (above the noise floor) precipitation inside of the plasmasphere.
- 2. If the event had been longer, we expect that the hiss wave would have also caused observed electron precipitation.

12:00

Thus we see how multiple loss mechanisms in this simply complicated event have occurred simultaneously, and affected each other.

12:00

3

NASA/Goddard Space Flight Center This has implications for more complicated events where other processes are likely also active and may further interact and either help or impede additional loss processes.