

Formation Process of the Outer Radiation Belt Through Nonlinear Interaction with Chorus Emissions Localized in Longitude

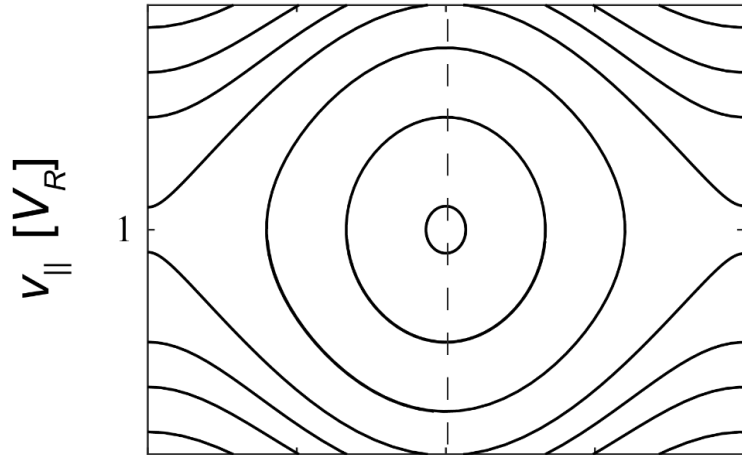
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Kyoto University

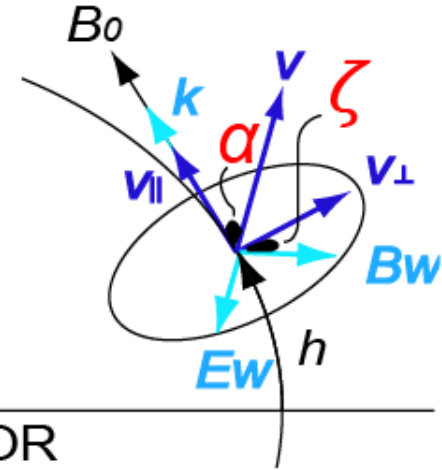
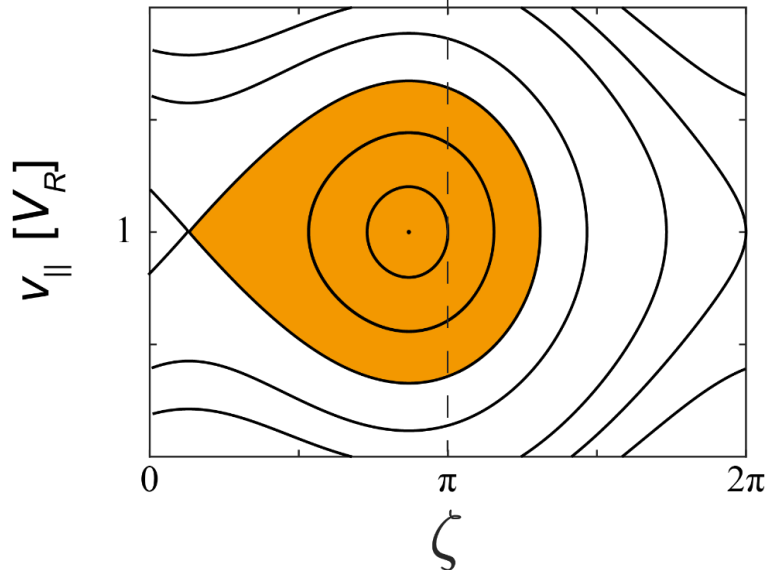
Electron Acceleration by Chorus Emissions

$$\mathbf{v} \cdot \mathbf{E}_w < 0 \quad \mathbf{v} \cdot \mathbf{E}_w > 0$$

(a) $S = 0$



(b) $-1 < S < 0$



EQUATOR

$$\frac{d\zeta}{dt} = \theta$$

$$\frac{d\theta}{dt} = \omega_{tr}^2 (\sin \zeta + S)$$

$$\theta = k(v_{\parallel} - V_R)$$

$$S = -\frac{1}{s_0 \omega \Omega_w} \left(s_1 \frac{\partial \omega}{\partial t} + c s_2 \frac{\partial \Omega_{ce}}{\partial h} \right)$$

[Omura et al., *JGR*, 2007, 2008] ²

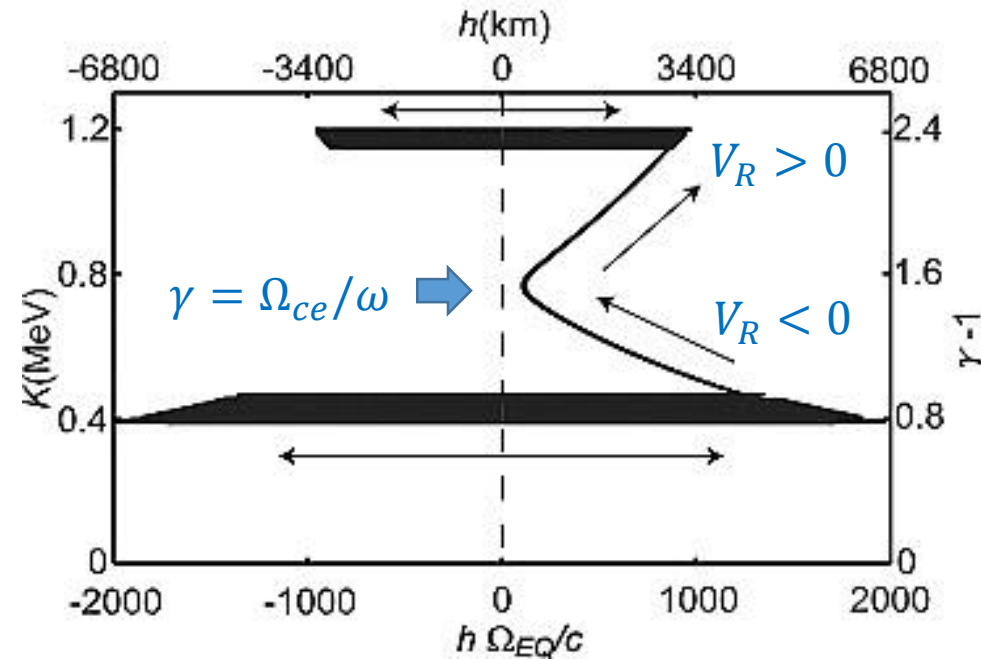
Two Nonlinear Acceleration Processes

$$V_R = \frac{\omega - \Omega_{ce}/\gamma}{k}$$

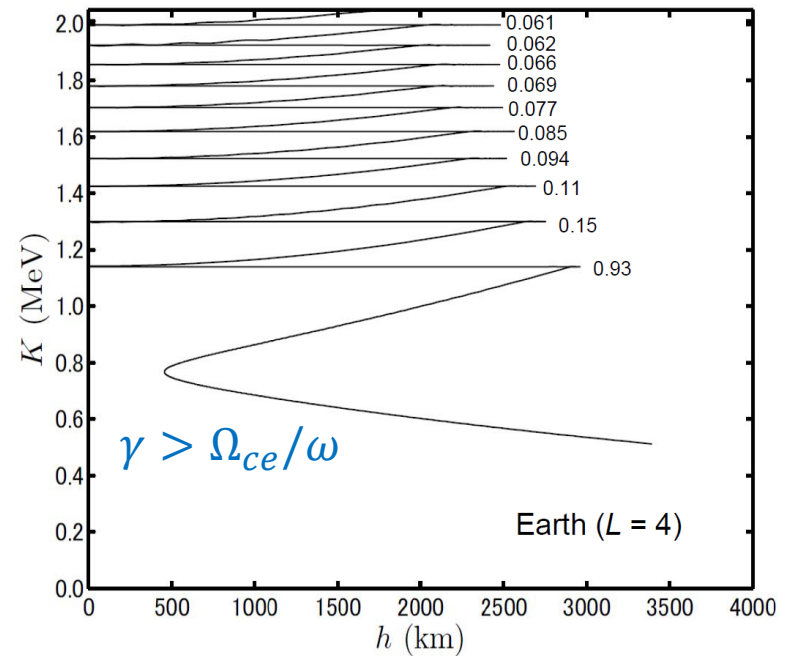
$$\gamma = \frac{1}{\sqrt{1 - (v/c)^2}}$$

Relativistic Turning Acceleration (RTA)

Ultra-Relativistic Acceleration (URA)

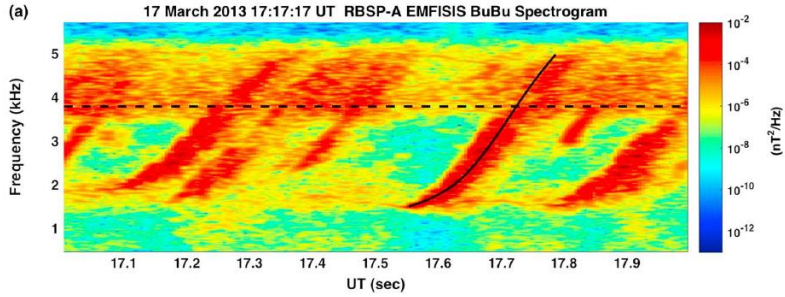


[Omura, et al., 2007]

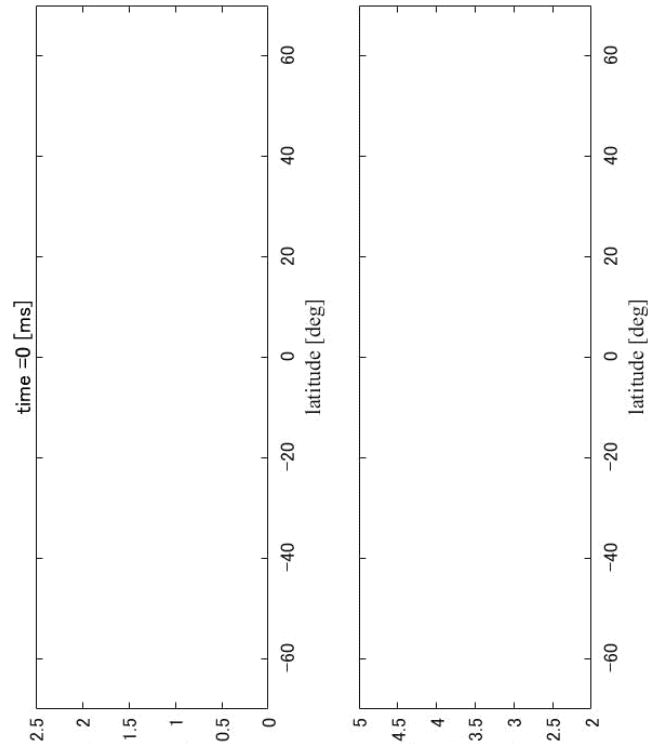
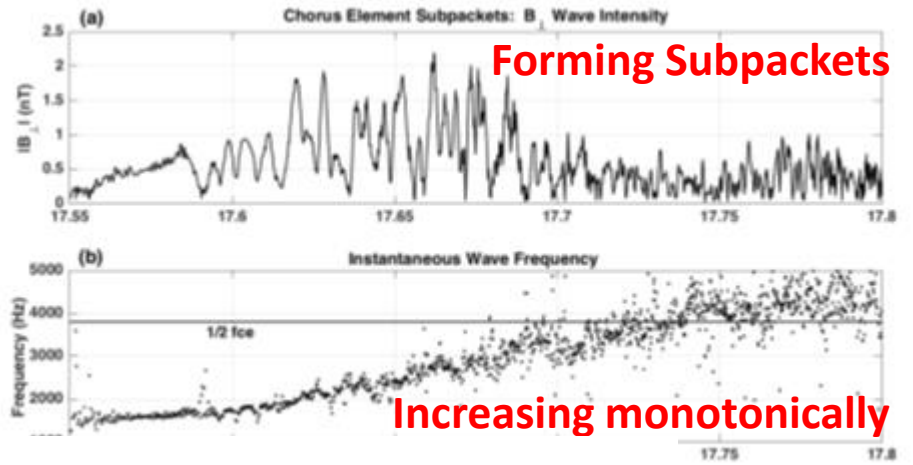
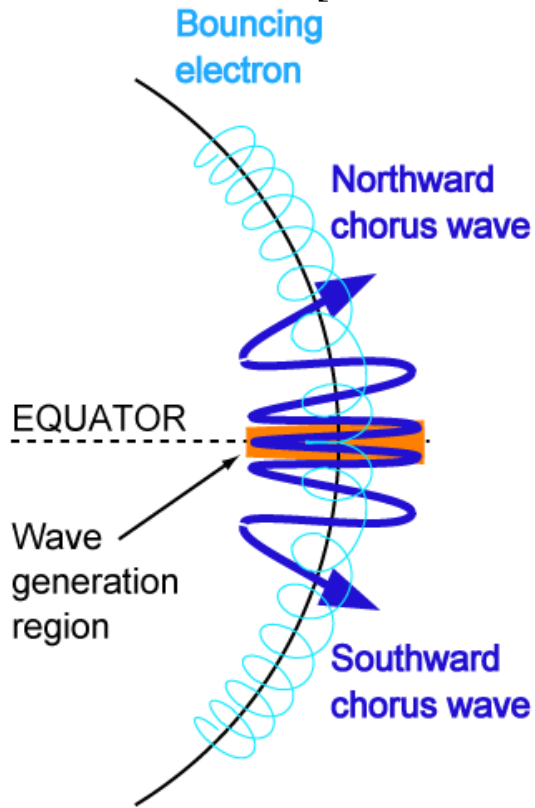


[Summers and Omura, 2007]

Chorus Waves



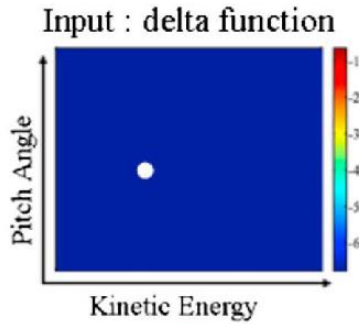
[Foster et al., 2017]



B_w [nT]

f [kHz]

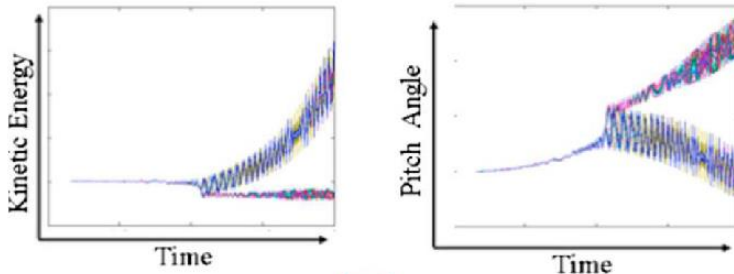
Numerical Green's Functions



$$\delta(E - E_0, \alpha - \alpha_0)$$

×

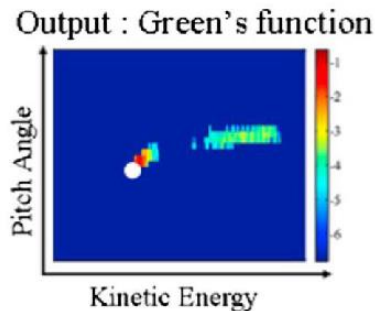
Test particle simulation through 1-cycle chorus emission



C

$$m_e \frac{d(\gamma \mathbf{v})}{dt} = -e [\mathbf{E}_w + \mathbf{v} \times (\mathbf{B}_0 + \mathbf{B}_w)]$$

||

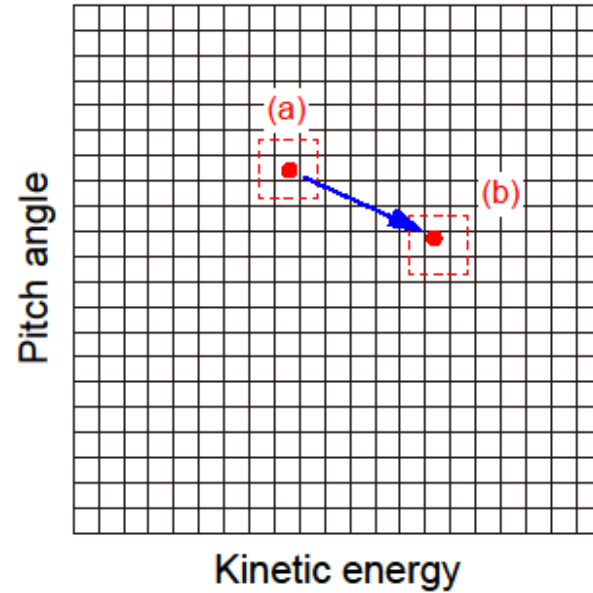


$$G(E, E_0, \alpha, \alpha_0)$$

Numerical Green's Functions

$$G(E, E_0, \alpha, \alpha_0) = \frac{\tilde{A}(E_0, \alpha_0)}{\tilde{A}(E, \alpha)} \times \frac{N_p(E, E_0, \alpha, \alpha_0)}{N_{p0}(E_0, \alpha_0) \Delta E \Delta \alpha}$$

$$N_{p0}(E_i, \alpha_j) = \sum_{E_k} \sum_{\alpha_l} N_p(E_k, E_i, \alpha_l, \alpha_j) .$$

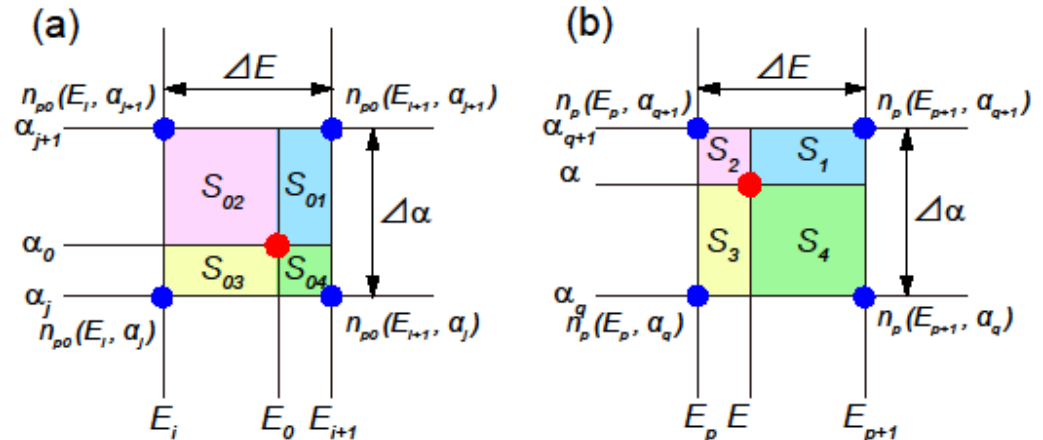


$$\int_{\alpha} \int_E \frac{N_p(E, E_0, \alpha, \alpha_0)}{N_{p0}(E_0, \alpha_0) \Delta E \Delta \alpha} dE d\alpha$$

$$= \sum_{\alpha} \sum_E \frac{N_p(E, E_0, \alpha, \alpha_0)}{N_{p0}(E_0, \alpha_0) \Delta E \Delta \alpha} \Delta E \Delta \alpha = 1$$



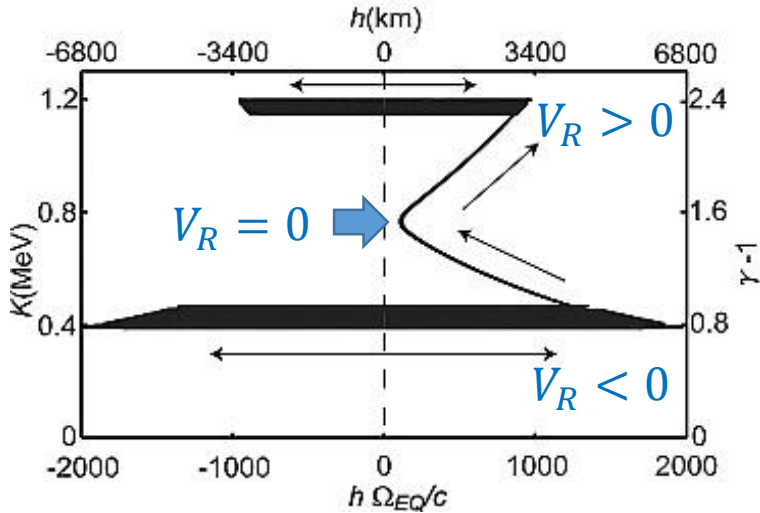
$$\int_{\alpha} \int_E \delta(E - E_0, \alpha - \alpha_0) dE d\alpha = 1$$



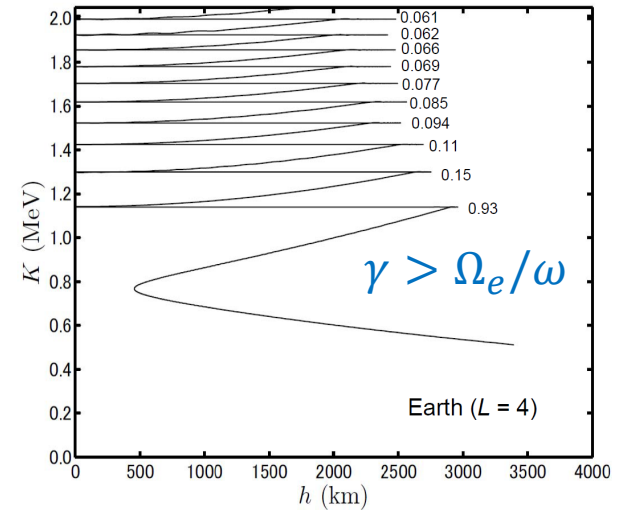
Relativistic Turning Acceleration (RTA)

Ultra-Relativistic Acceleration (URA)

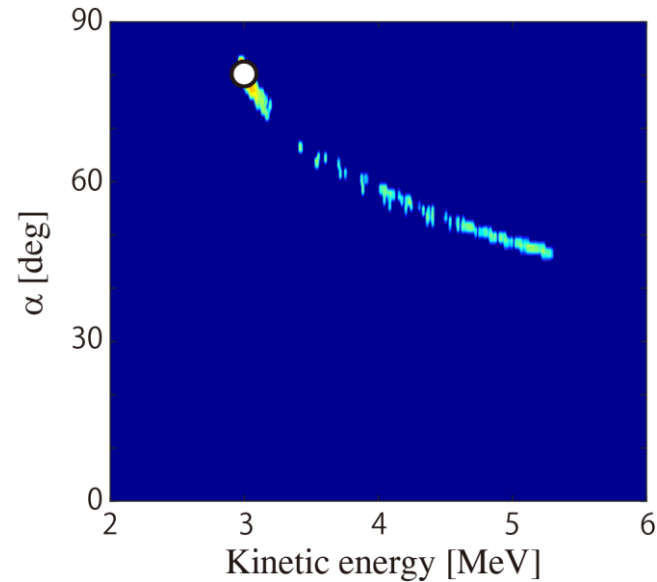
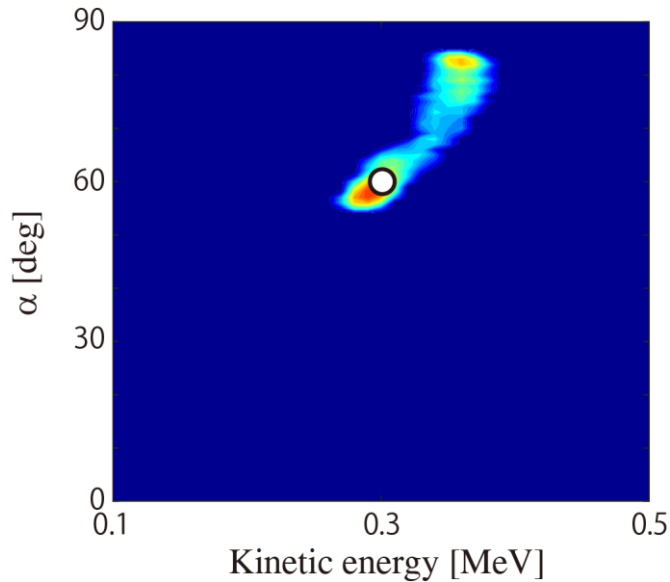
$$V_R = \frac{\omega - \Omega_{ce}/\gamma}{k}$$



[Omura, et al, JGR, 2007]

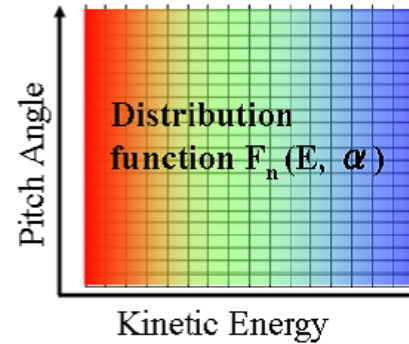


[Summers and Omura, GRL, 2007]



Numerical Green's Function Method

Input : Electron distribution function at n cycles

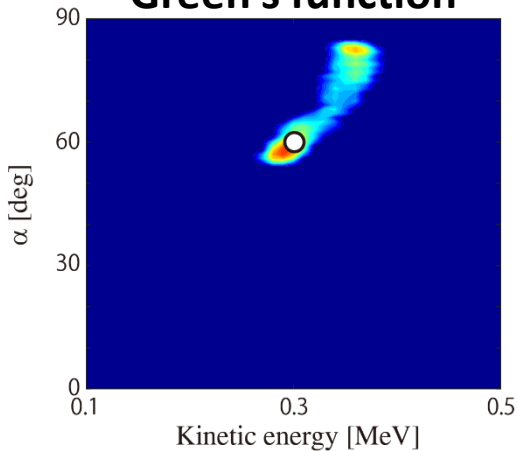


Test particle simulation: **Linearity**

$$C[f_a + f_b] = C[f_a] + C[f_b]$$

$$WC[f_a] = C[wf_a]$$

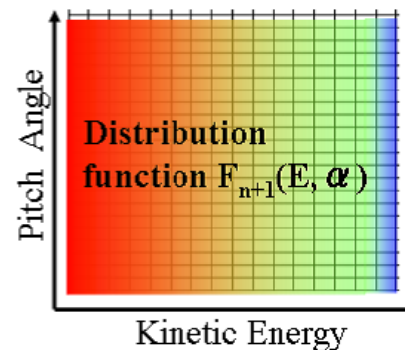
Green's function



Convolution integral

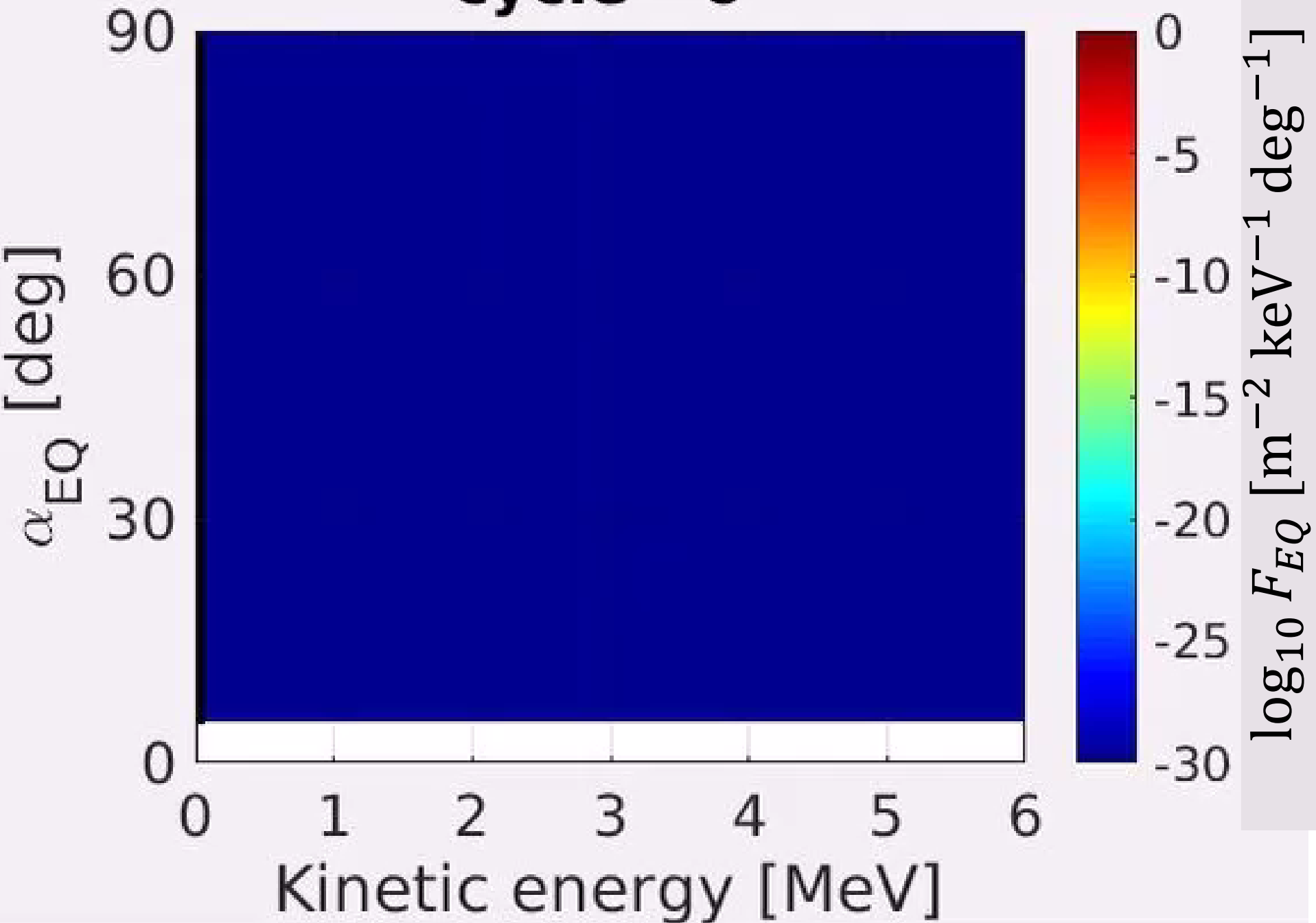
$$f_n(E, \alpha) = \sum_{\alpha_0} \sum_{E_0} f_{n-1}(E_0, \alpha_0) G(E, E_0, \alpha, \alpha_0) \Delta E \Delta \alpha$$

Output : Distribution after n+1 cycle chorus emission



[Omura et al., 2015]

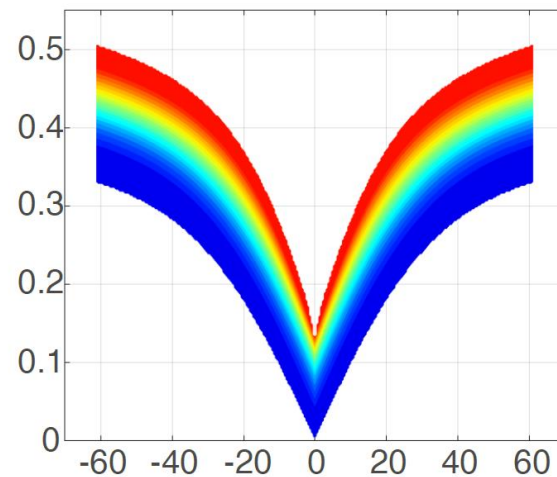
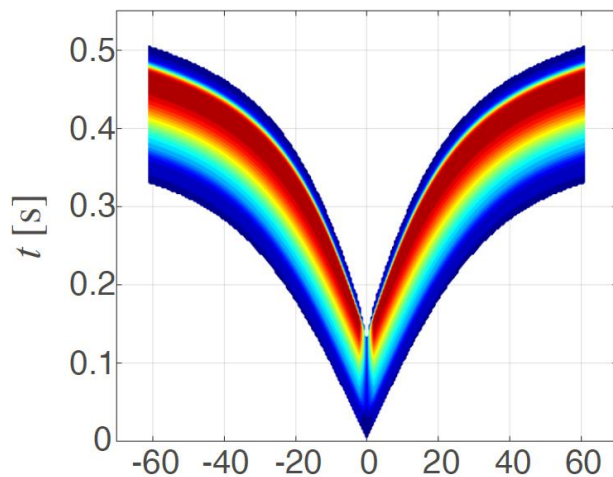
cycle = 0



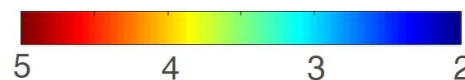
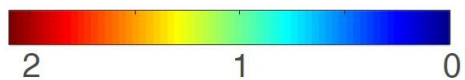
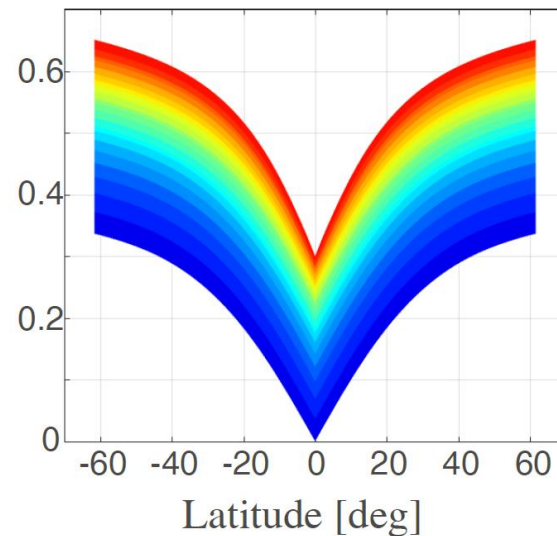
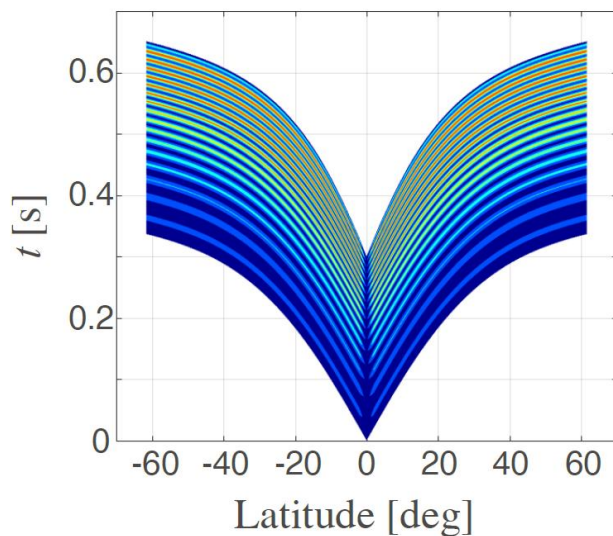
Wave amplitude [nT]

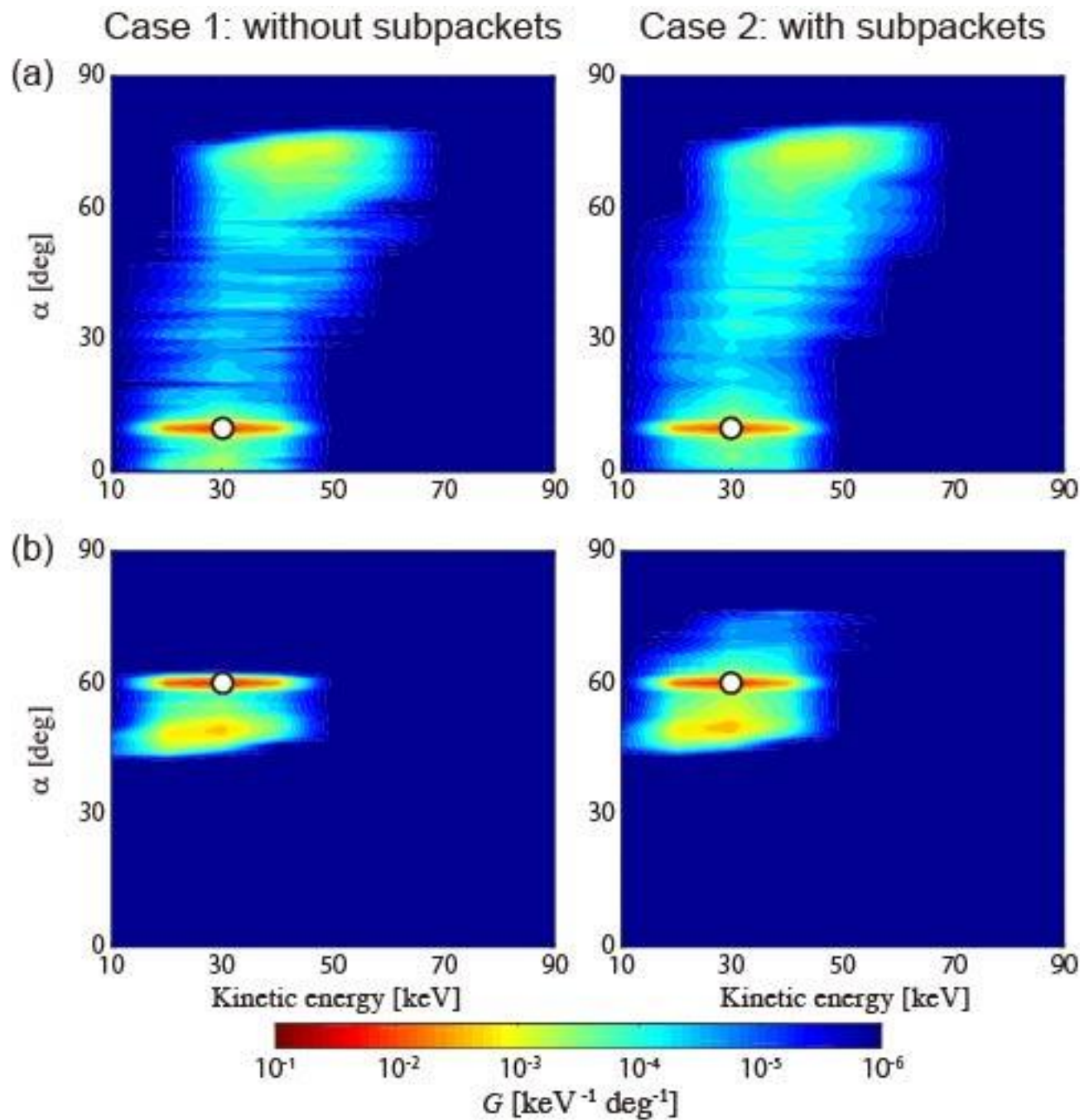
Wave frequency [kHz]

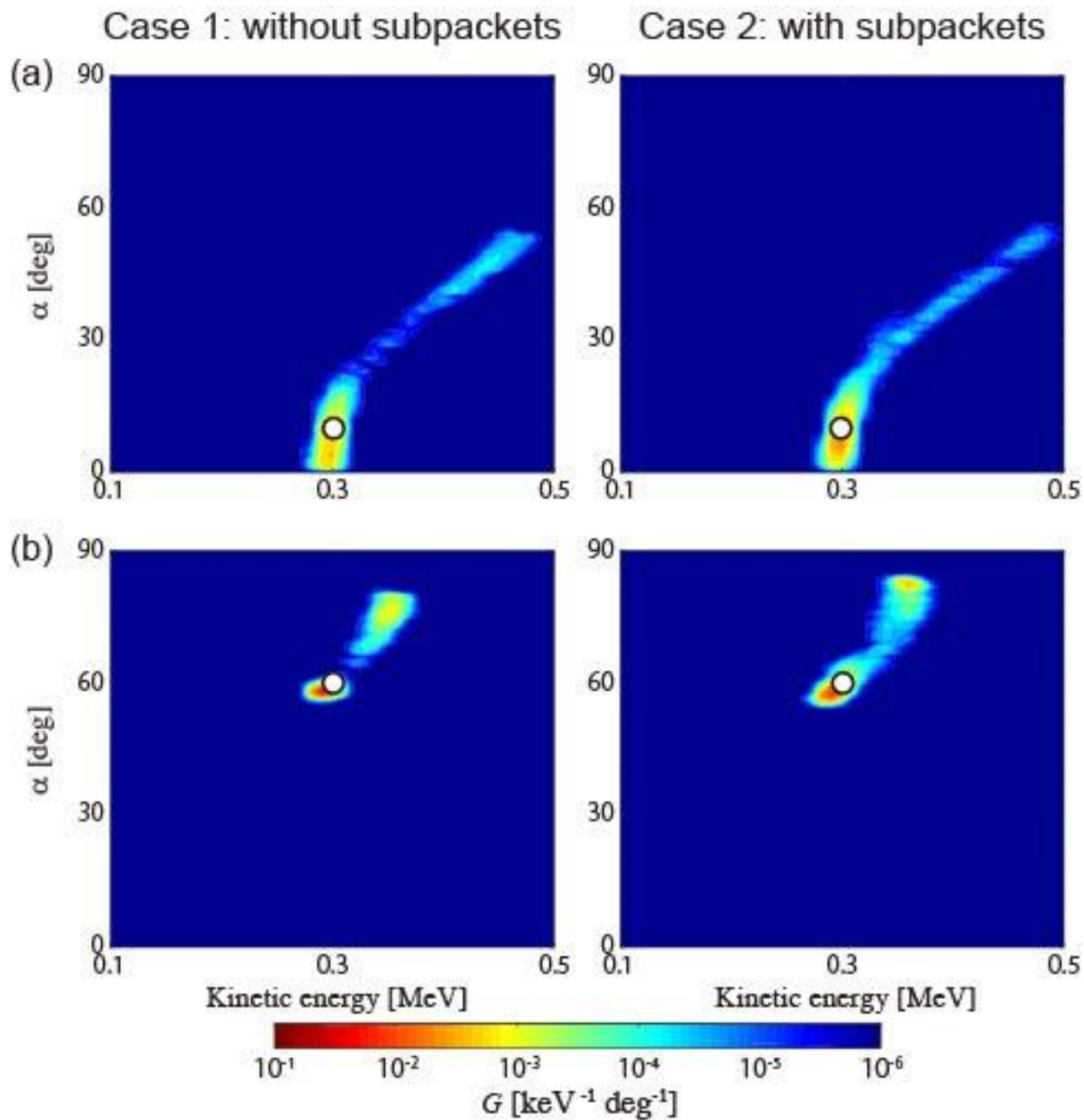
(a) Case 1: without subpackets

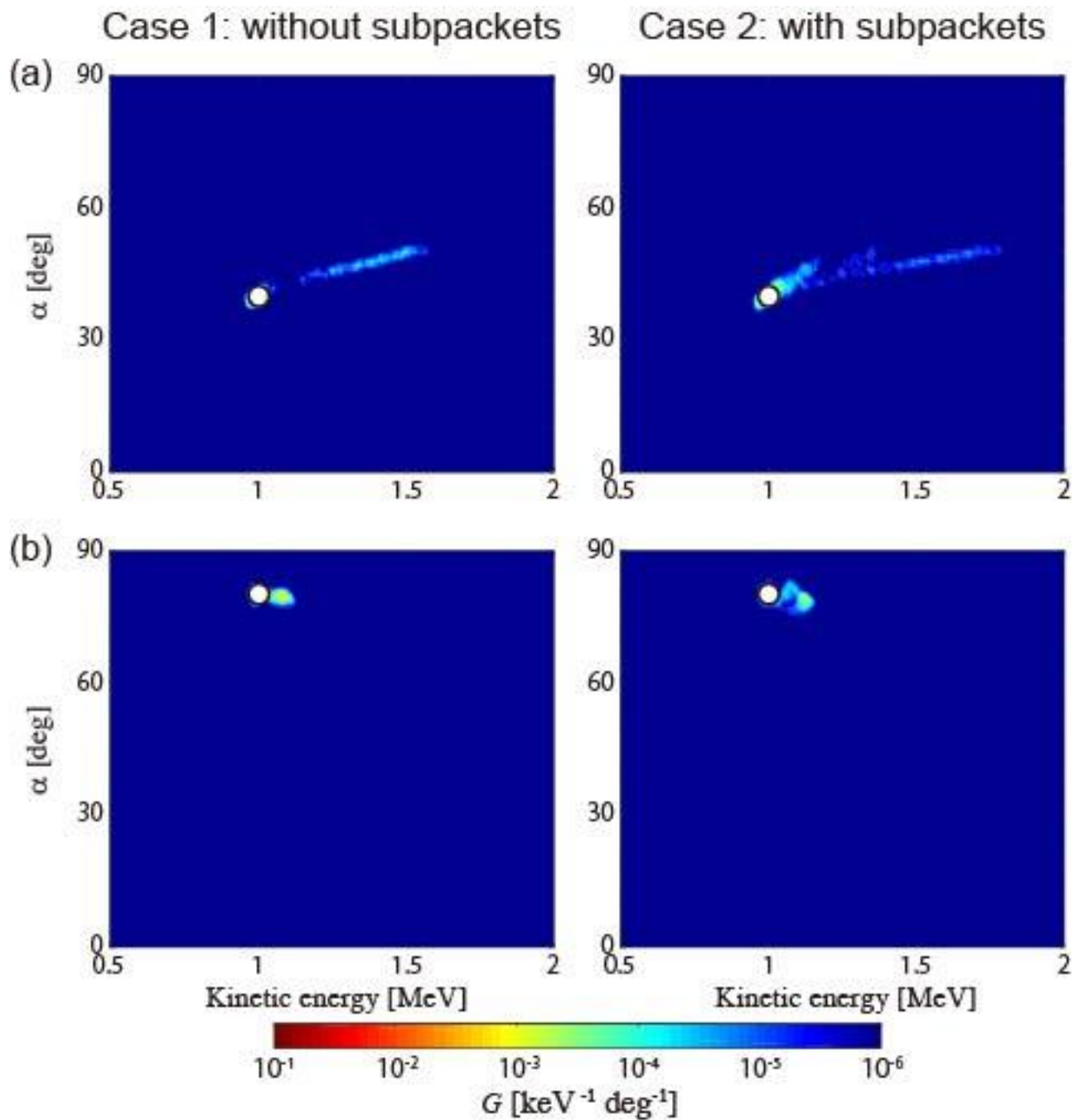


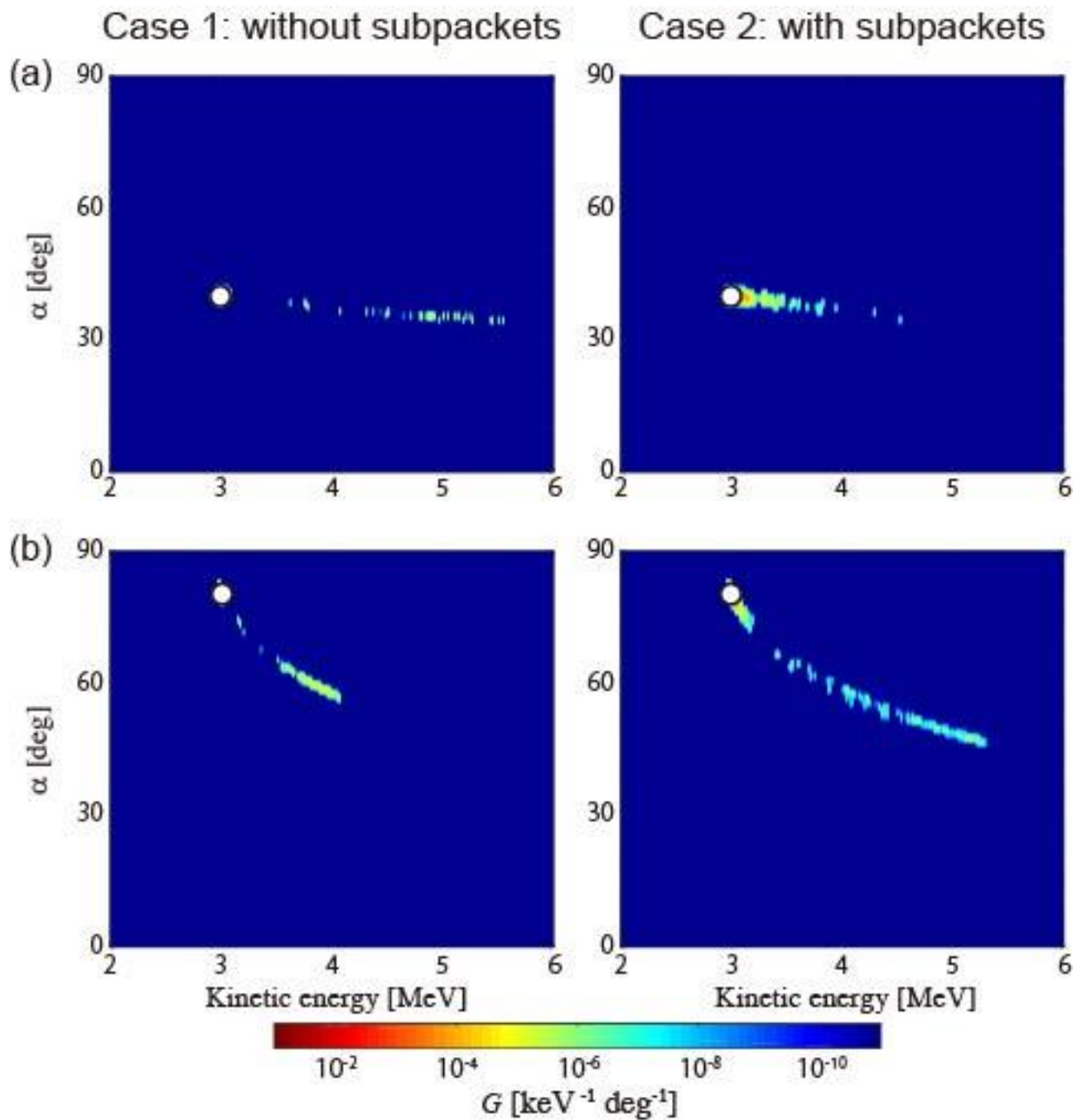
(b) Case 2: with subpackets







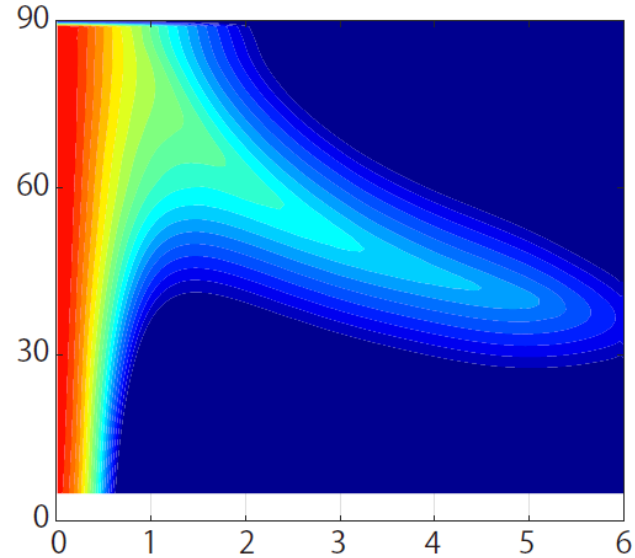
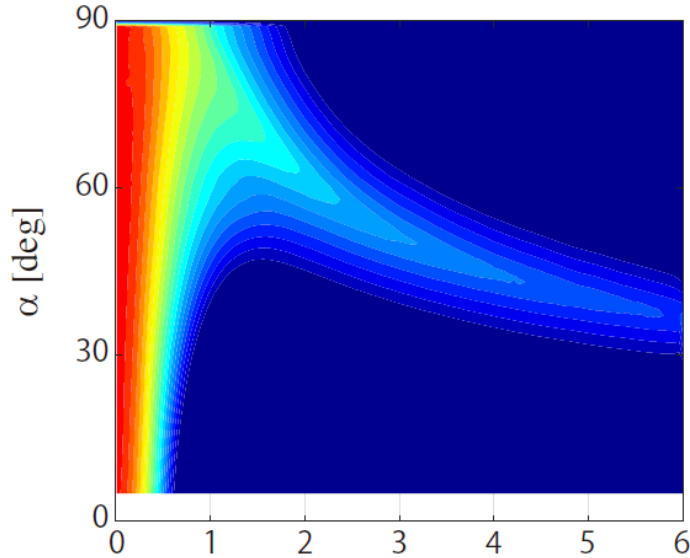




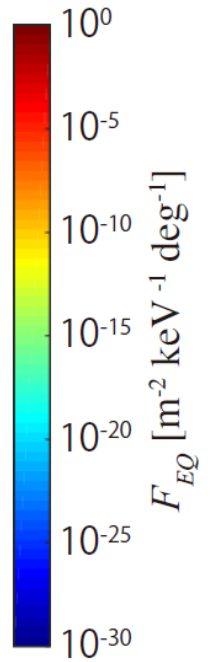
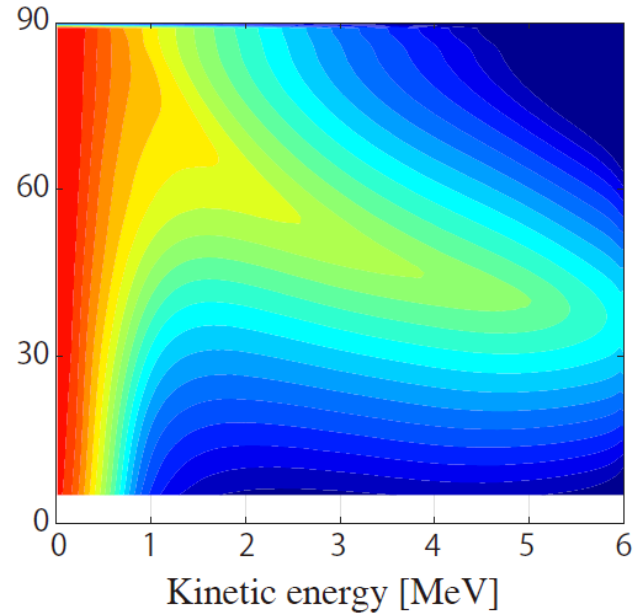
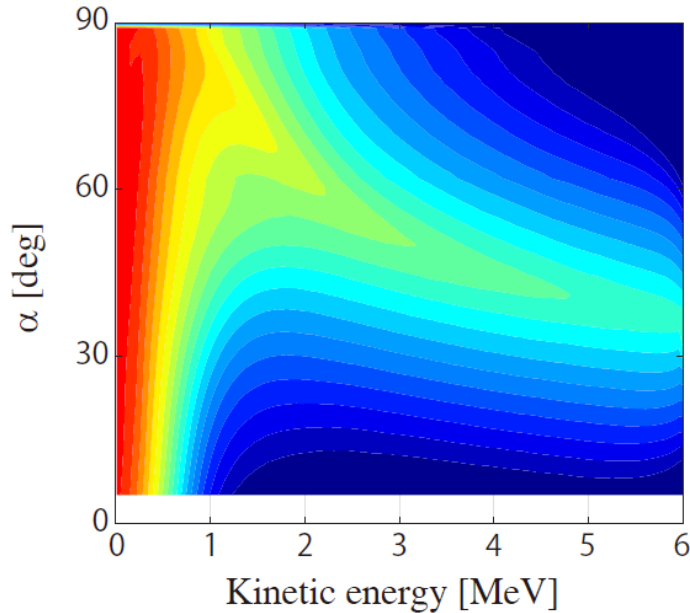
Case 1: without subpackets

Case 2: with subpackets

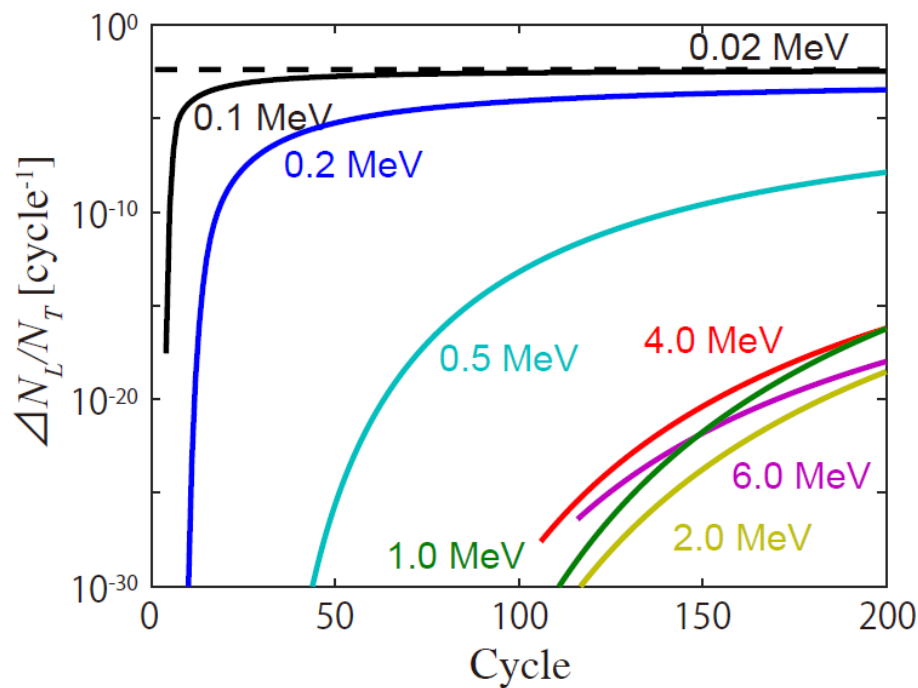
(a) cycles = 100



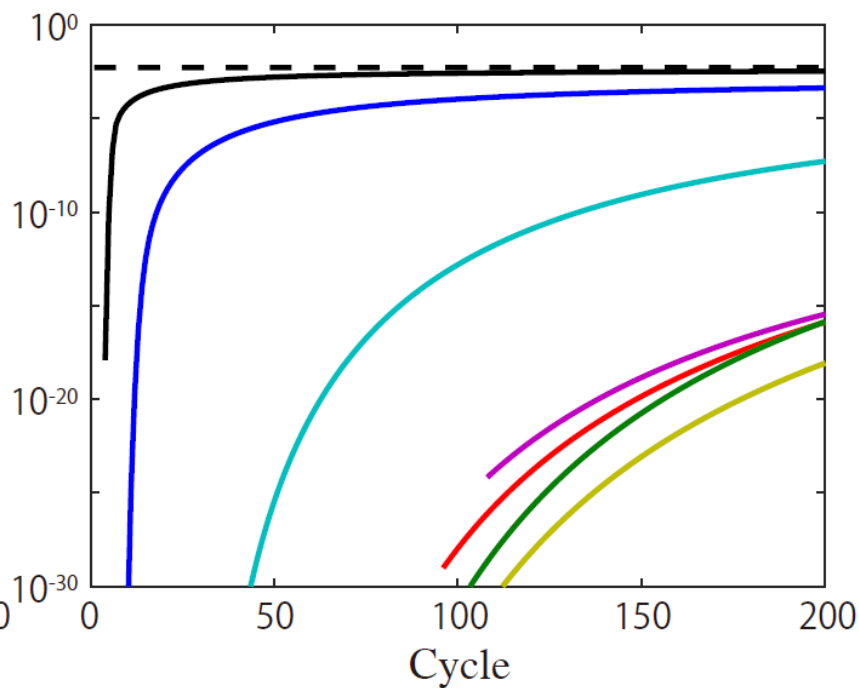
(b) cycles = 200



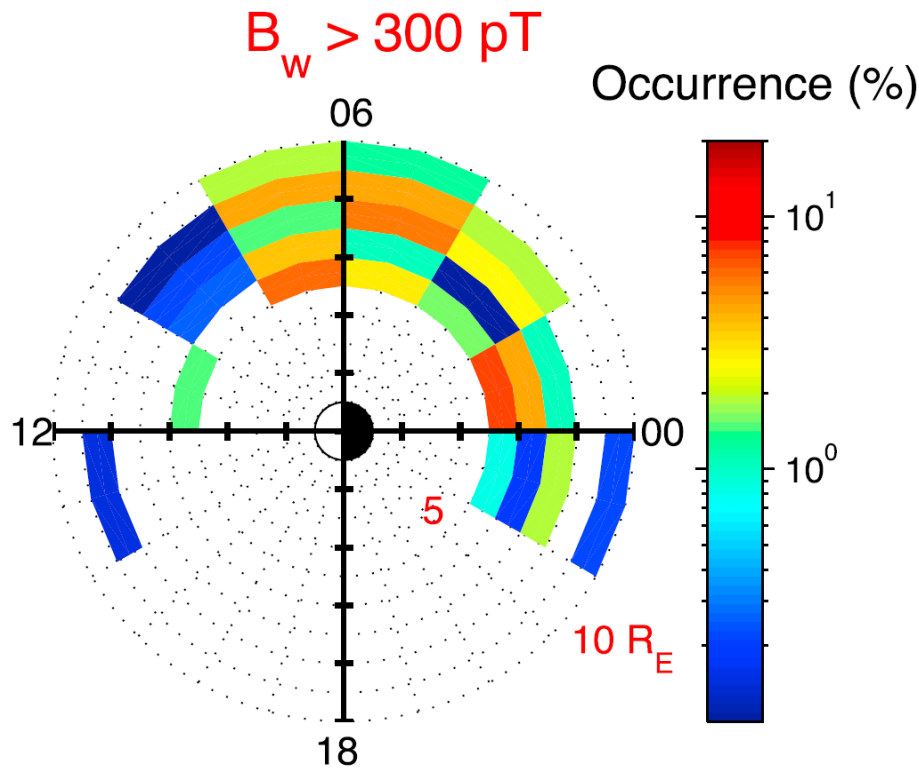
Case 1: without subpackets



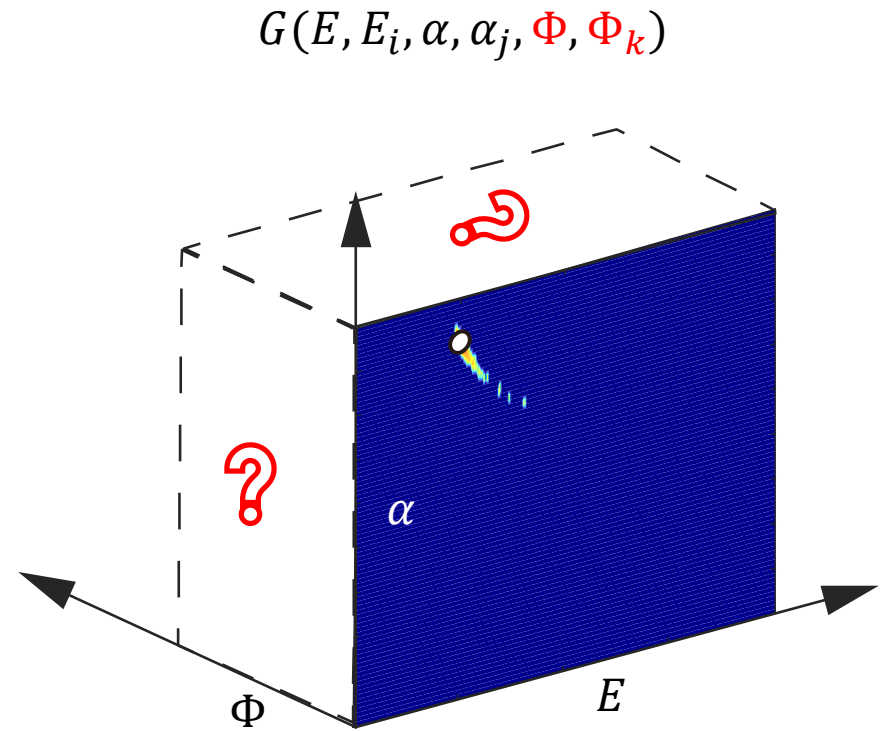
Case 2: with subpackets



Longitudinal Dependency



[Li et al., JGR, 2011]



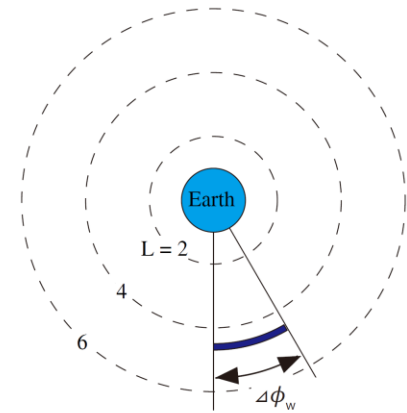
Modified Numerical Green's Function Method

✧ Assuming dipole geomagnetic field

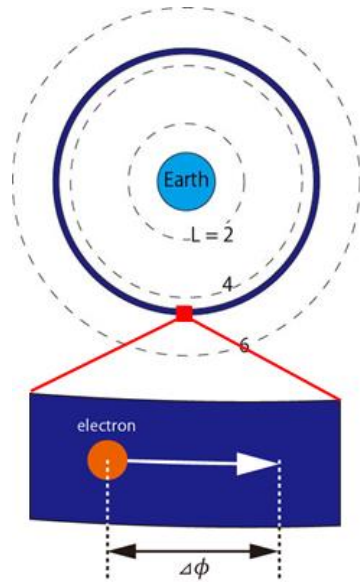
✧ $\Delta\Phi \leq 3^\circ$ at $E \leq 6$ MeV and $\alpha \leq 90^\circ$

$$f_n^\phi(E, \alpha, \phi) = \sum_{a_0} \sum_{E_0} \sum_{\phi_w} f_{n-1}^\phi(E_0, \alpha_0, \phi_w) G^\phi(E, E_0, \alpha, \alpha_0, \phi - \phi_w) \Delta\phi \Delta E \Delta\alpha$$

$$+ \sum_{\bar{\phi}_w} f_{n-1}^\phi(E, \alpha, \bar{\phi}_w) G_0^\phi(E, \alpha, \phi - \bar{\phi}_w) \Delta\phi,$$

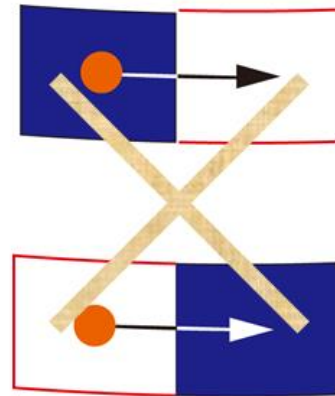


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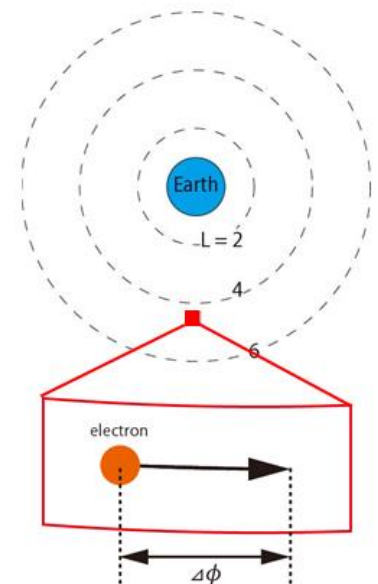
$G^\phi(E, E_0, \alpha, \alpha_0, \Delta\phi_d)$

+



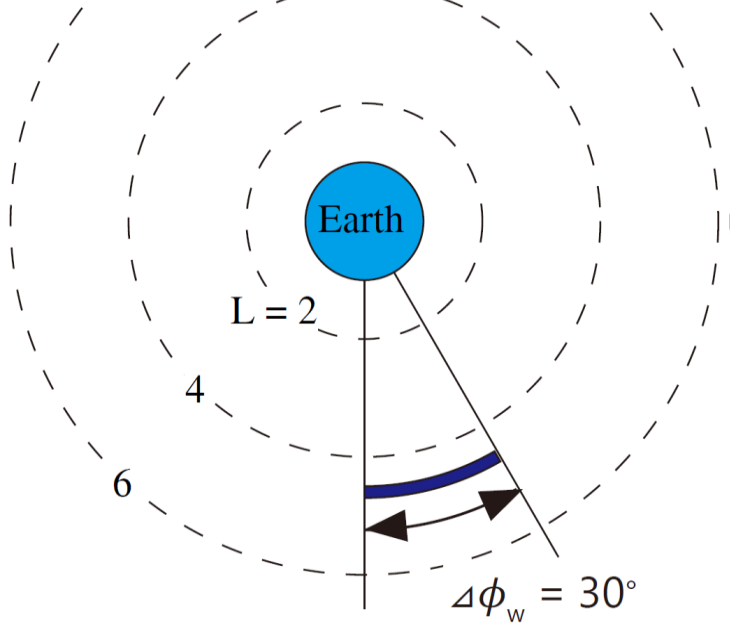
Neglected

+

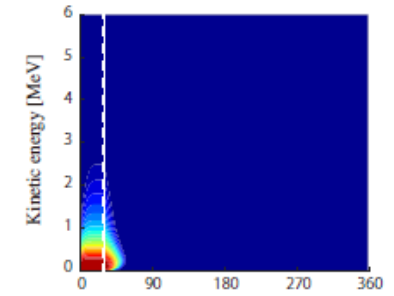
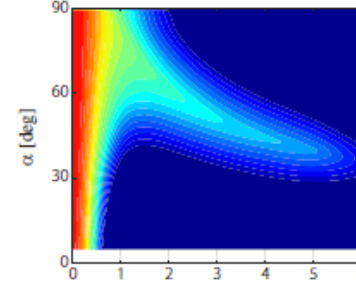


$G_0^\phi(E, \alpha, \Delta\phi_d)$

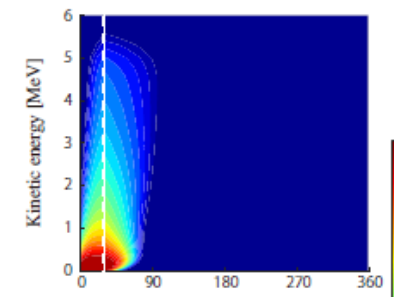
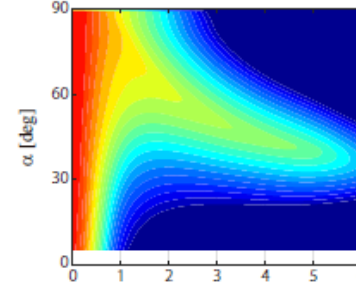
Constant influx: 10 – 30 keV electrons



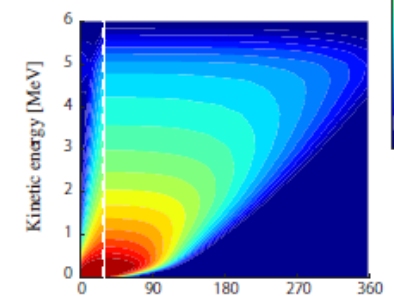
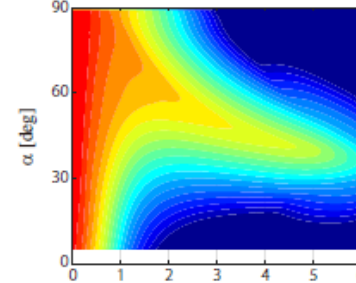
(a) cycles = 100, $t = 65.2$ s



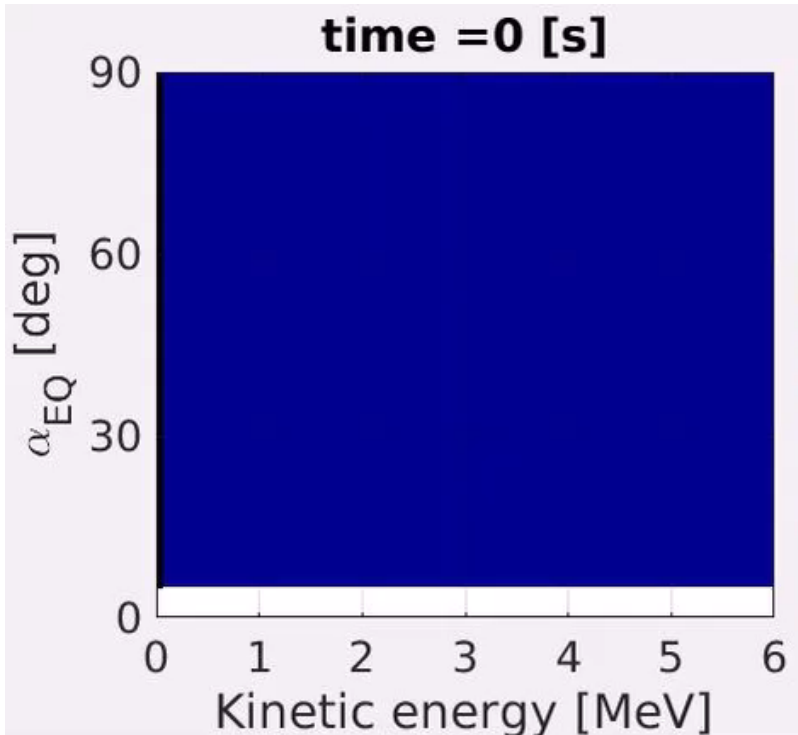
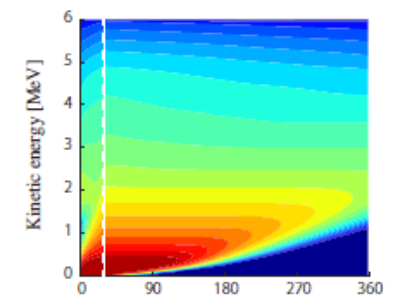
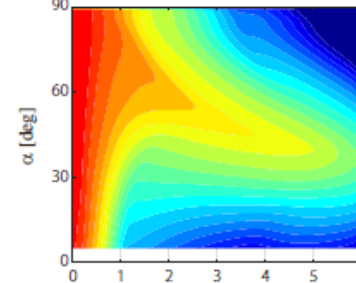
(b) cycles = 200, $t = 130$ s



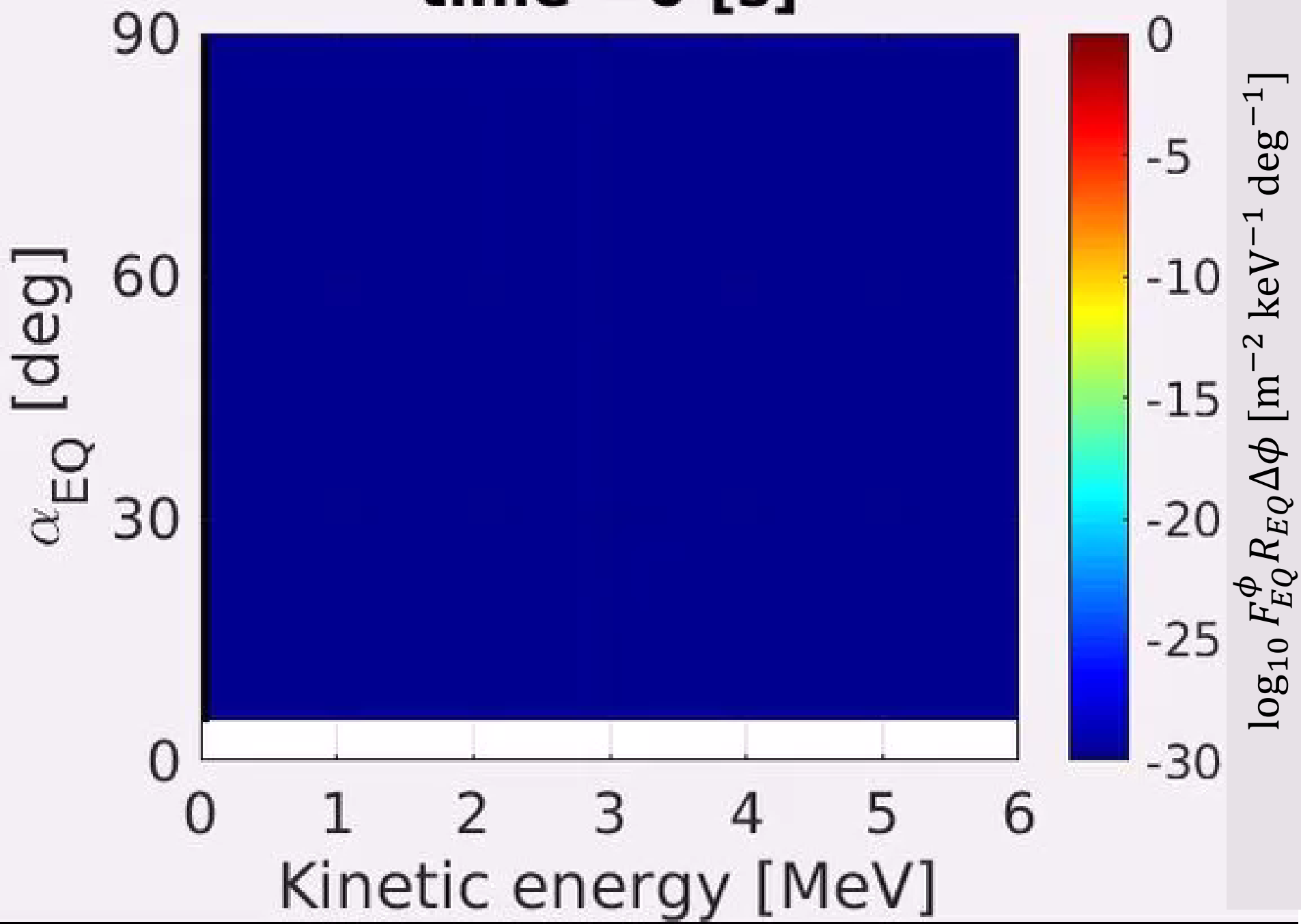
(c) cycles = 500, $t = 326$ s



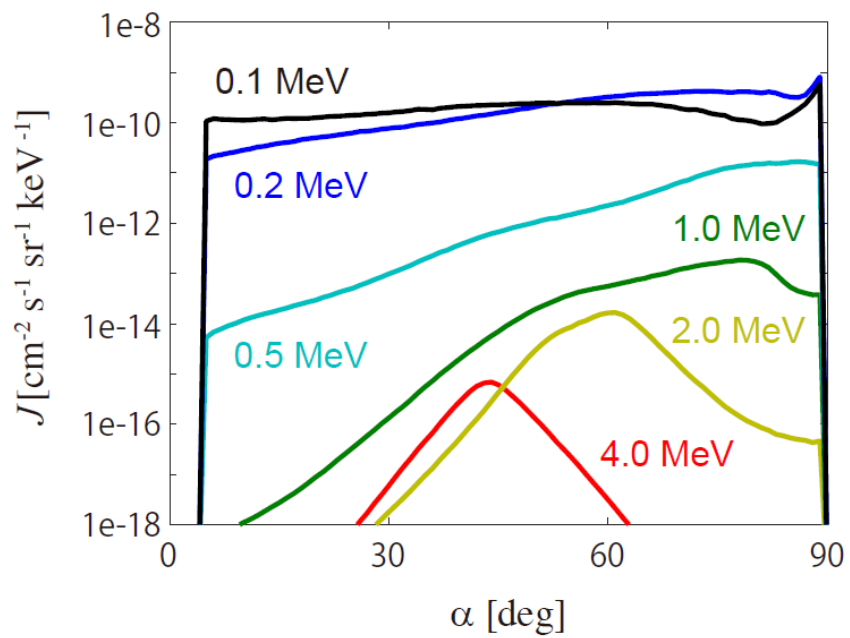
(d) cycles = 1000, $t = 652$ s



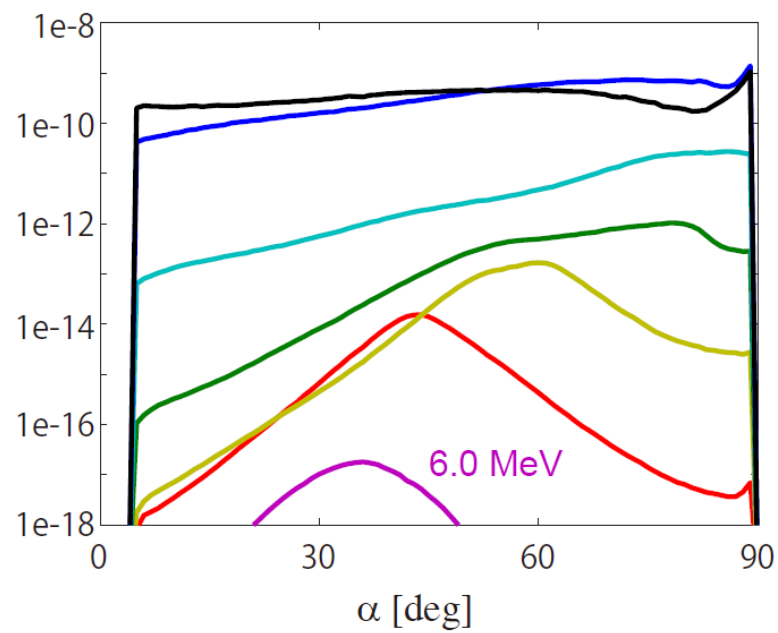
time = 0 [s]



(a) $t = 0.54$ hour (3000 cycles)



(b) $t = 1.09$ hour (6000 cycles)

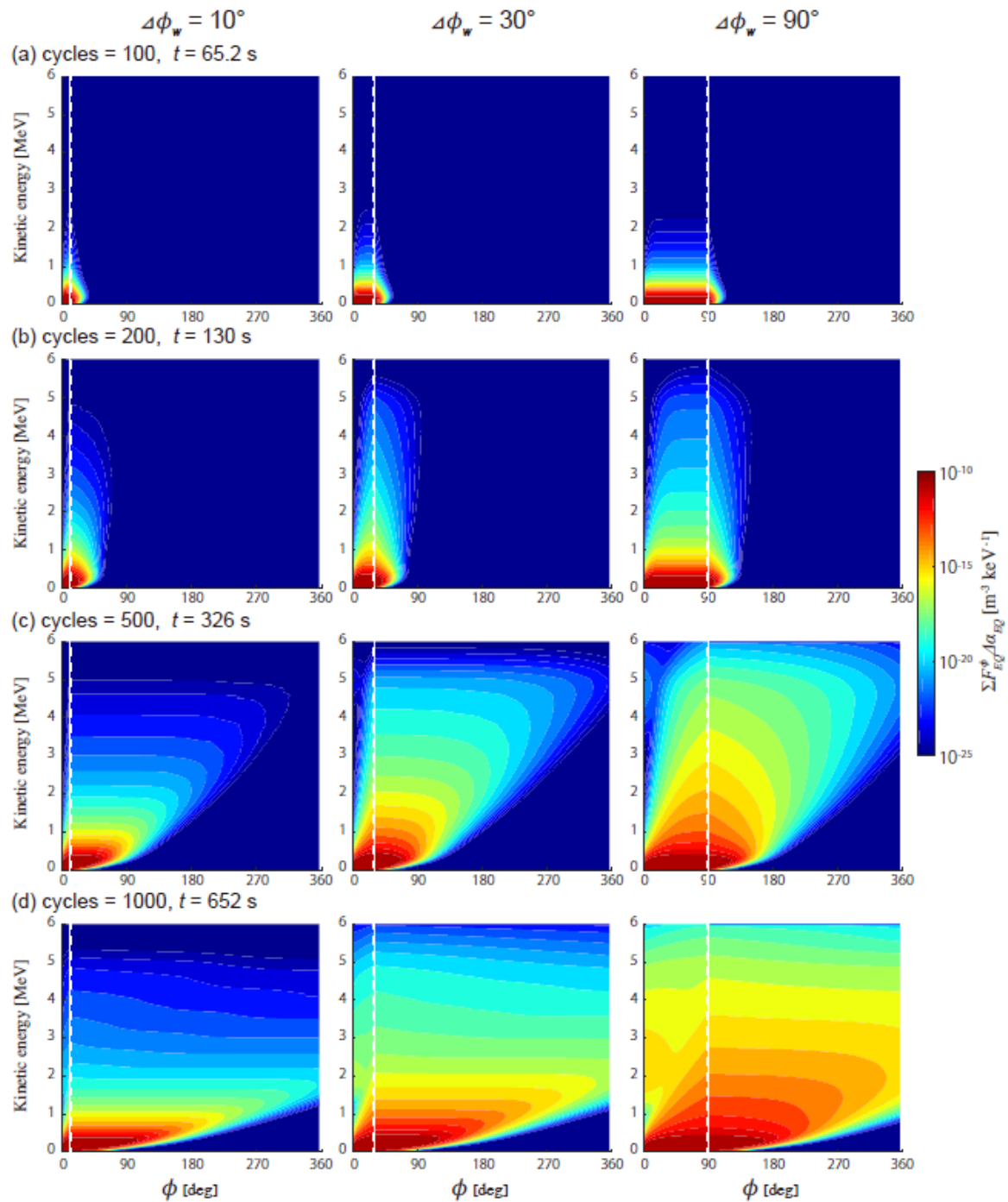


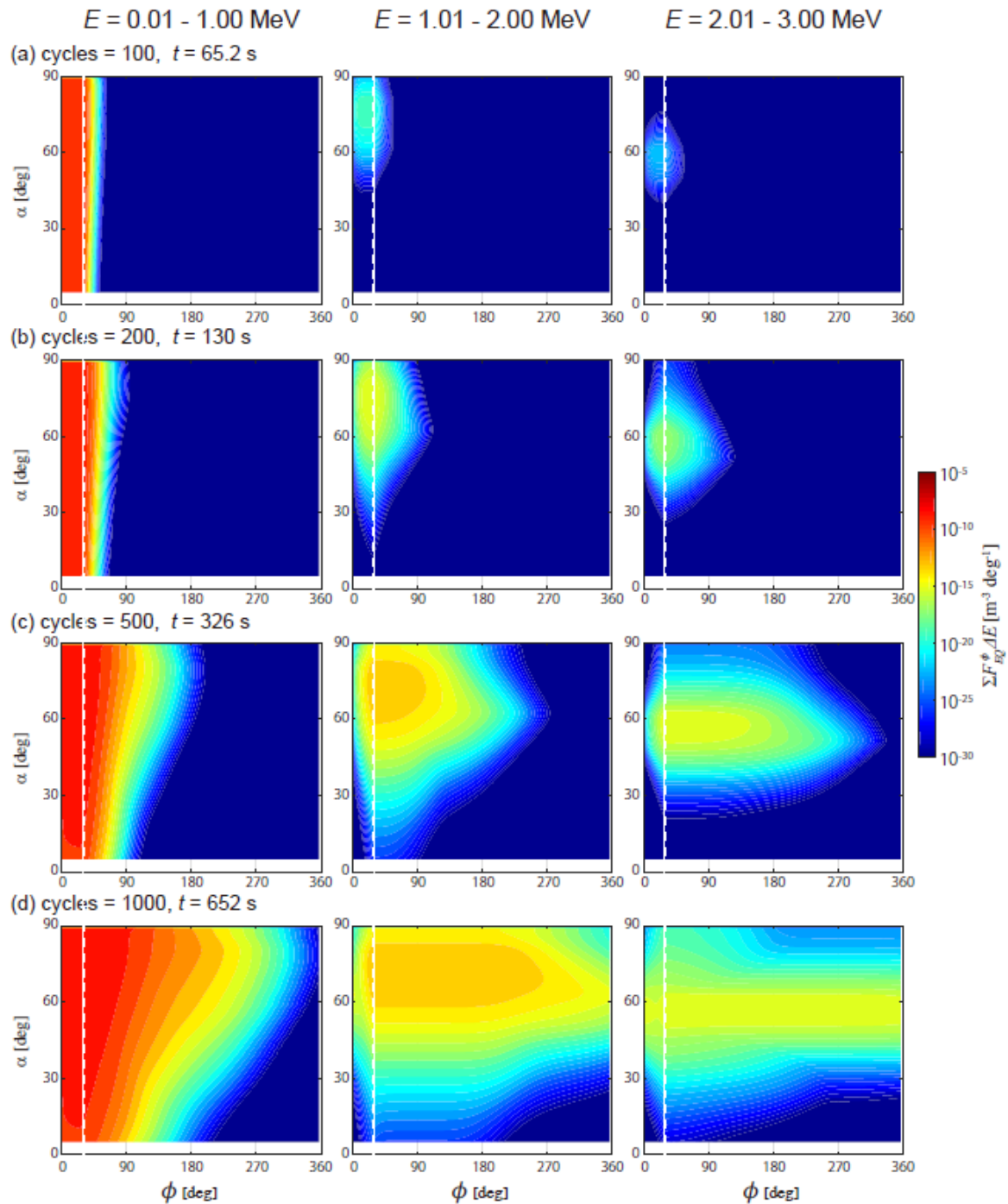
Summary

1. We have modified the numerical Green's function method including **longitudinal dependency** of chorus emissions.
2. Chorus emissions **localized in longitude** result in rapid formation of the outer radiation belt.

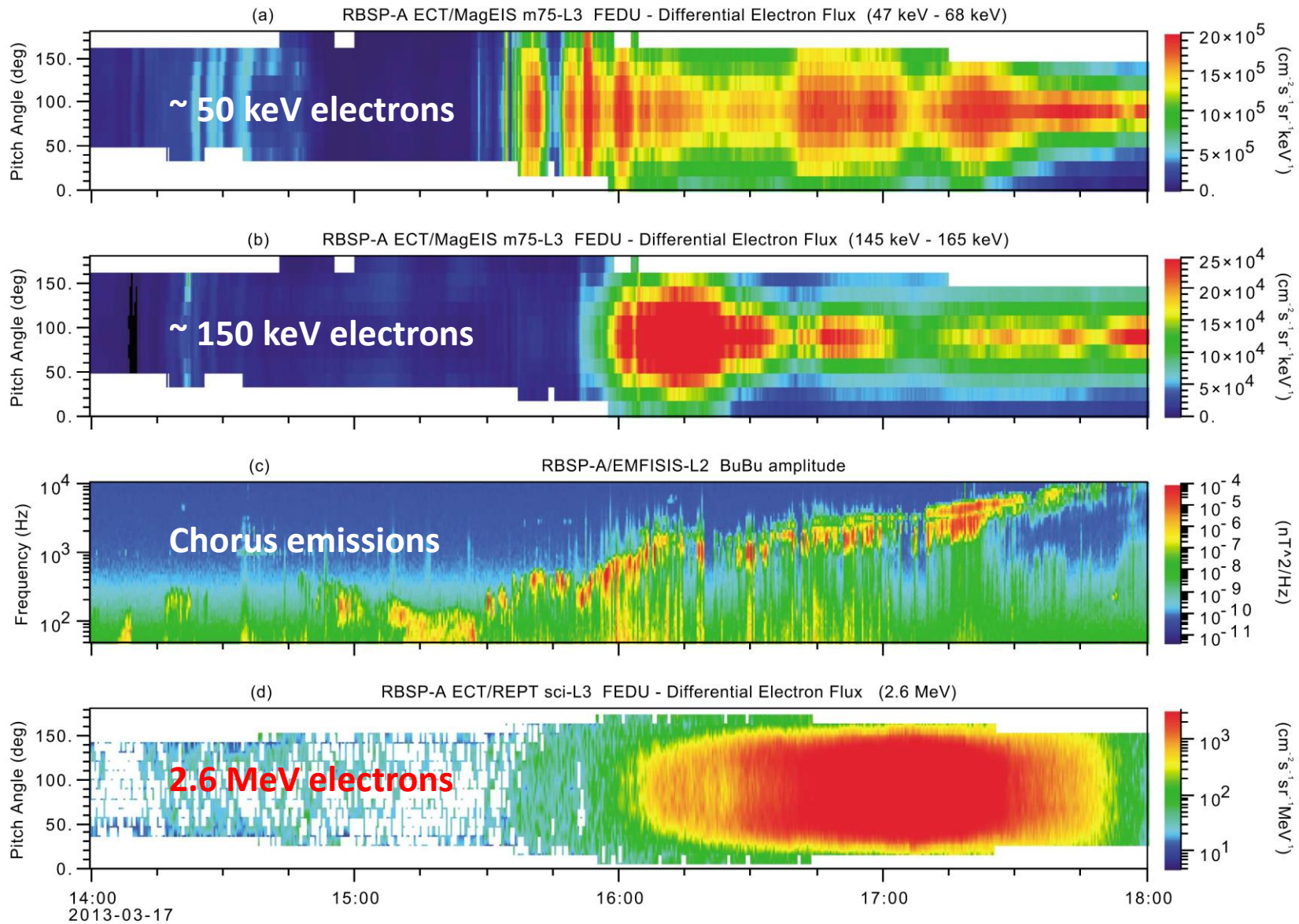
References

1. Kubota, Y., and Y. Omura, Nonlinear dynamics of radiation belt electrons interacting with chorus emissions localized in longitude, submitted to *J. Geophys. Res.*
2. Omura, Y. et al., Formation process of relativistic electron flux through interaction with chorus emissions in the Earth's inner magnetosphere, *J Geophys. Res. Space Physics*, 120, 9545-9562, 2015.





Rapid-acceleration of MeV Electrons



[Foster et al., 2017]